



# ANR NEWS (LDM) - R&D R2D2 (2 $\beta$ ) (and the SPC)

Conseil Scientifique IN2P3 - Octobre 2018

Pascal Lautridou pour les groupes IN2P3 de R2D2 & NEWS

# What approach



It is based on the advantages provided by the Spherical Proportional Counter (SPC)

Material budget & Volume/Surface => background radioactivity

For large mass detectors => modular systems ... but cheap

Gaseous medium => gas type & pressure

Signal readout => signal processing

Detection performances => threshold & energy resolution

# Principle of the SPC

I. Giomataris JINST 2008

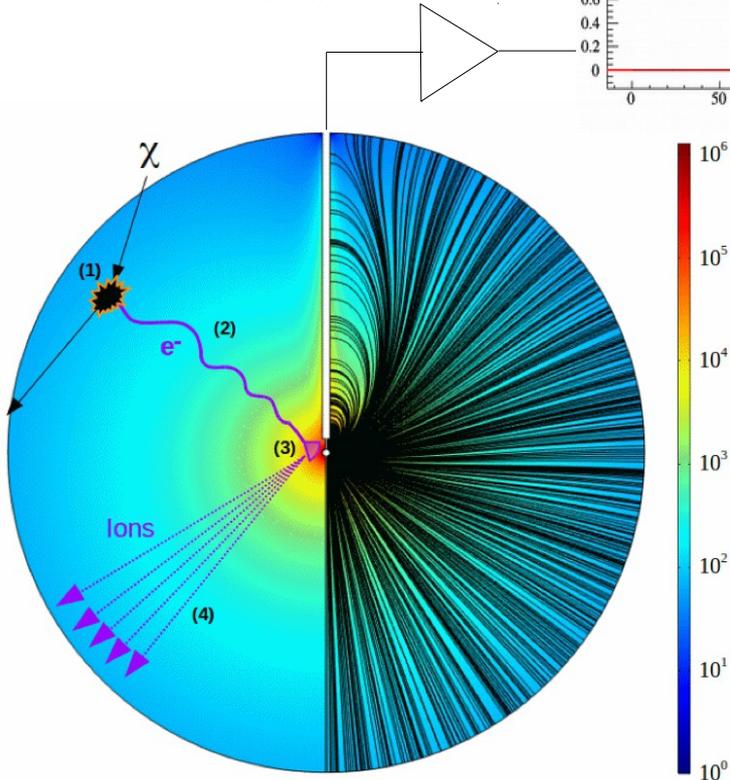
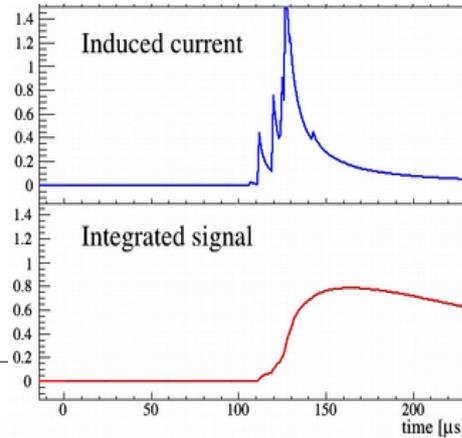
$$C = 4\pi\epsilon\rho < 1 \text{ pF}$$

$$1/\rho = 1/r_a - 1/r_c$$

$$\rho \approx r_a$$

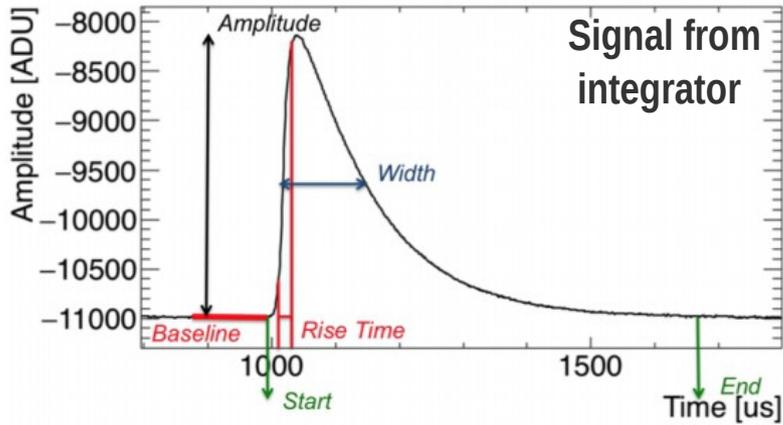
$$E(r) = V_0 r_a / r^2$$

P. Gros UCLA Dark Matter conference 2018



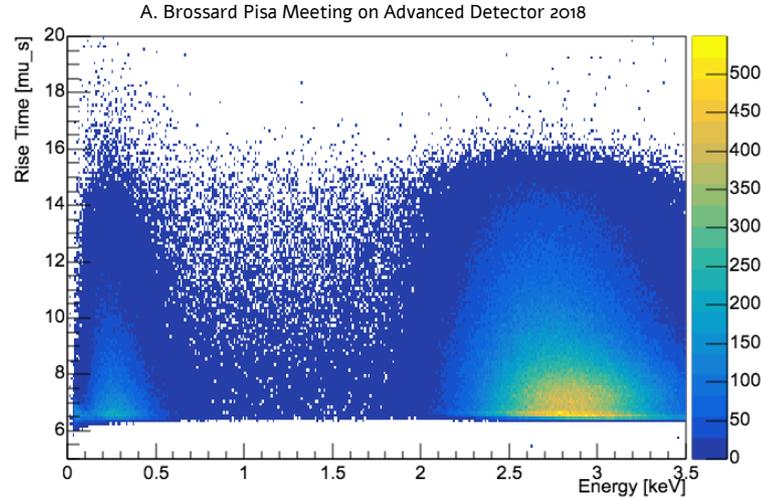
- Simple mechanical structure
- Easy adjustment of parameters
  - Sensor:  $\phi > 1 \text{ mm}$
  - Gain:  $< 10^4$
- Single channel FEE + MHz BWD integrator => allows signal processing
- Threshold down to single electron
- Identification of point like energy deposition via time dispersion:  $\sigma(r) \approx (r/r_c)^3$

# Energy Resolution & Energy Threshold

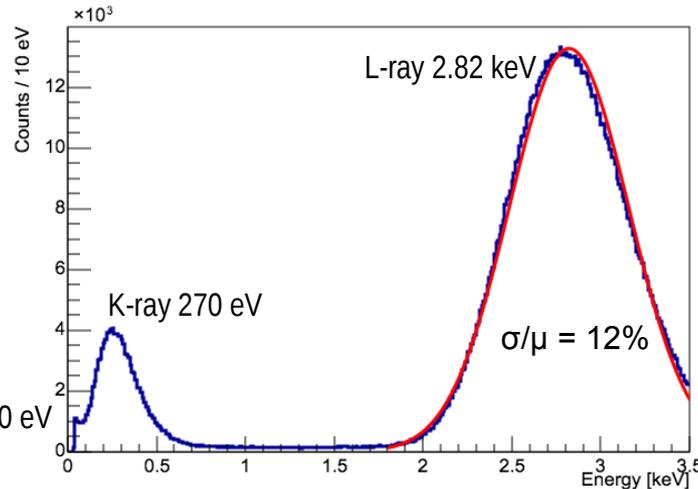


Rise Time ~longitudinal diffusion  
~ radius of energy deposition

Threshold < 50 eV



0.5 bar Ar/CH4 (2 %) + Ar 37 source



@ 1 MeV  
 $\Delta E/E < 1\%$   
FWHM  
expected

scaling suggests  
0.3 - 0.7 %

# ANR NEWS (New Experiments With Spheres) @ LSM

## Dark Matter search in 0.1 - 10 GeV

A. Dastgheibi-Fard, I. Giomataris\*, M. Gros, O. Guillaudin, I. Katsioulas, J.-F. Muraz, J.-P. Mols, X.-F. Navick, T. Papaevangelou, F. Piquemal, D. Santos, M. Zampaolo

*LSM, LPSC, IRFU + Queen's University, Canada (13 Members)\**

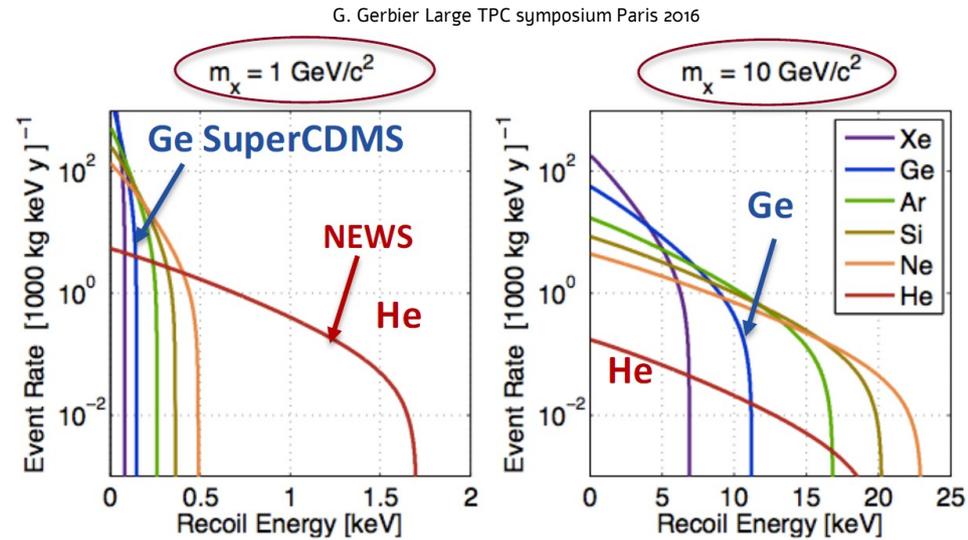
### **ANR Goal (2015-2019): development of the methodology**

As part of NEWS-G (New Experiments With Spheres-Gas)

*+ Pacific Northwest National Laboratory, USA, (2 Members) + Royal Military College of Canada, Canada, (2 Members) + University München, Germany, (1 Member) + SNOLAB, Canada, (1 Member) + Aristotle University of Thessaloniki, Greece, (1 Member) + University of Birmingham, UK, (1 Member) + University of Alberta, USA, (1 Member) + TRIUMF laboratory, Canada, (1 Member) + SUBATECH, (1 Member)*

# Light nuclei (H, He, Ne) to detect light WIMPs

Recoil distributions with various targets



**First phase: use SEDINE @ LSM as prototype for detection  
with He, Ne @ 3 bars**

**Tasks:** sphere construction, low-radioactivity material, cleaning methods, sensor & FEE optimization, DAQ set-up, calibration procedures, quenching factor

# 2015-2019: SEDINE @ LSM

## 2016: 3.1 bar Ne/CH<sub>4</sub> (0.7 %) in sealed mode

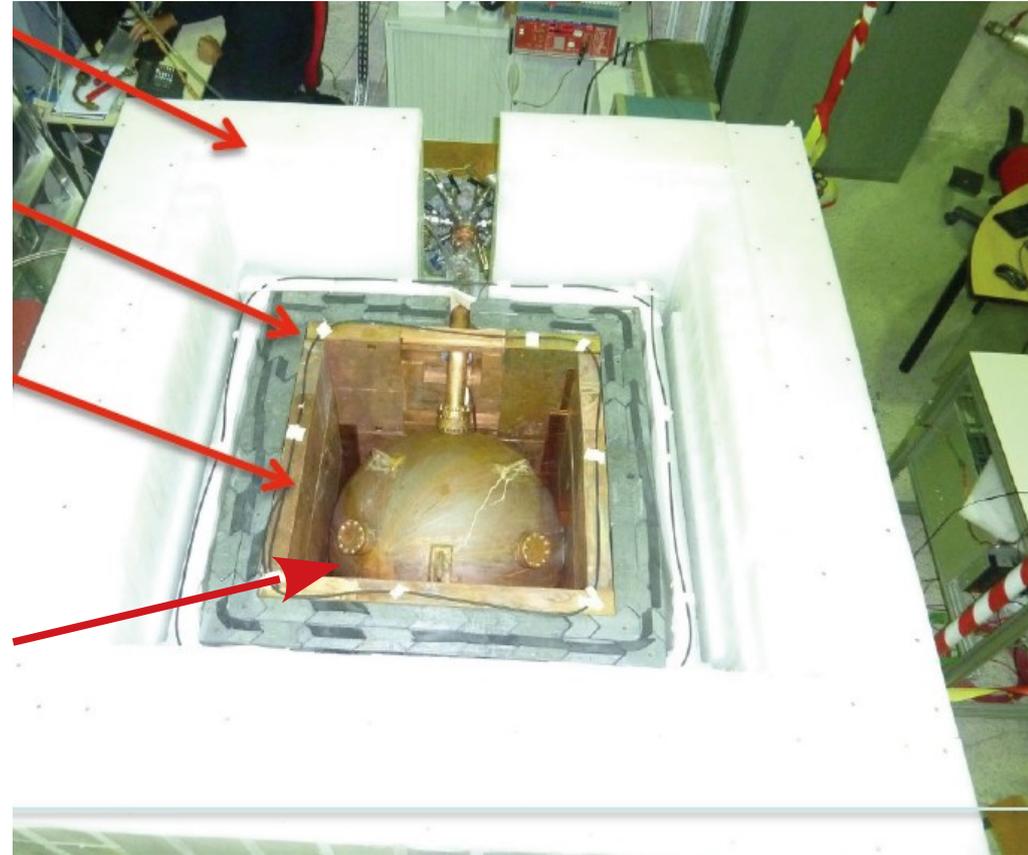
- Ionization energy  $w = 36$  eV
- HV: 2520 V, no sparks, **Gain: 3000**
- FEE: Canberra 2006 (50  $\mu$ s RC), Sampling: 2 MS/s
- Acquisition threshold @ 50 eV (~30 ADU)
- Data taking continuously during 42 days
- **Loss of gain 3 % along 42 days** monitored with
- 210 Po line + variation on days scale of  $\pm 4$  %
- Calibrations: 37Ar & 8 keV line from Cu fluorescence  
+ simulation Am-Be neutron source

30 cm PE

15 cm Pb

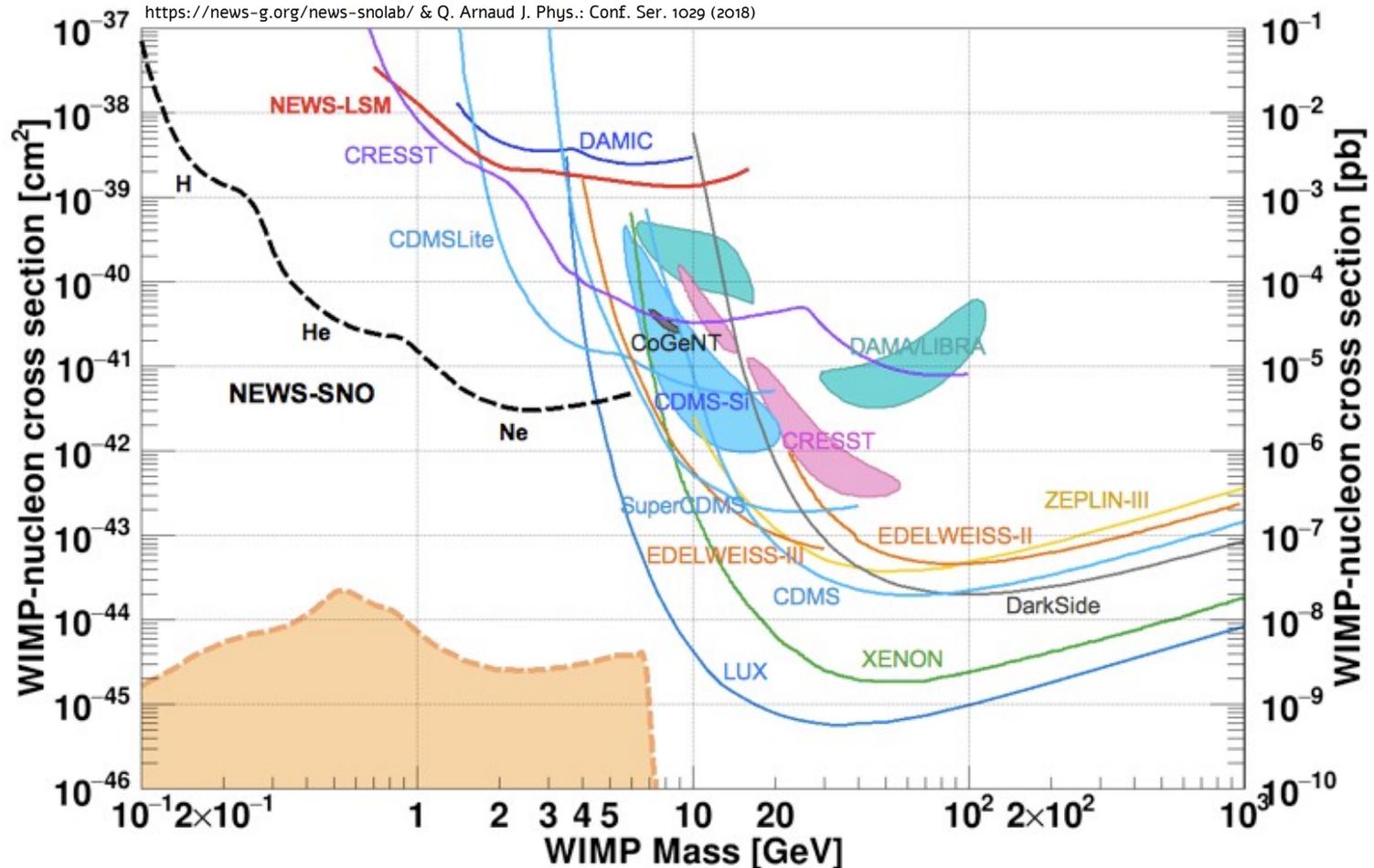
8 cm Cu

$\phi$  61 cm  
Sphere



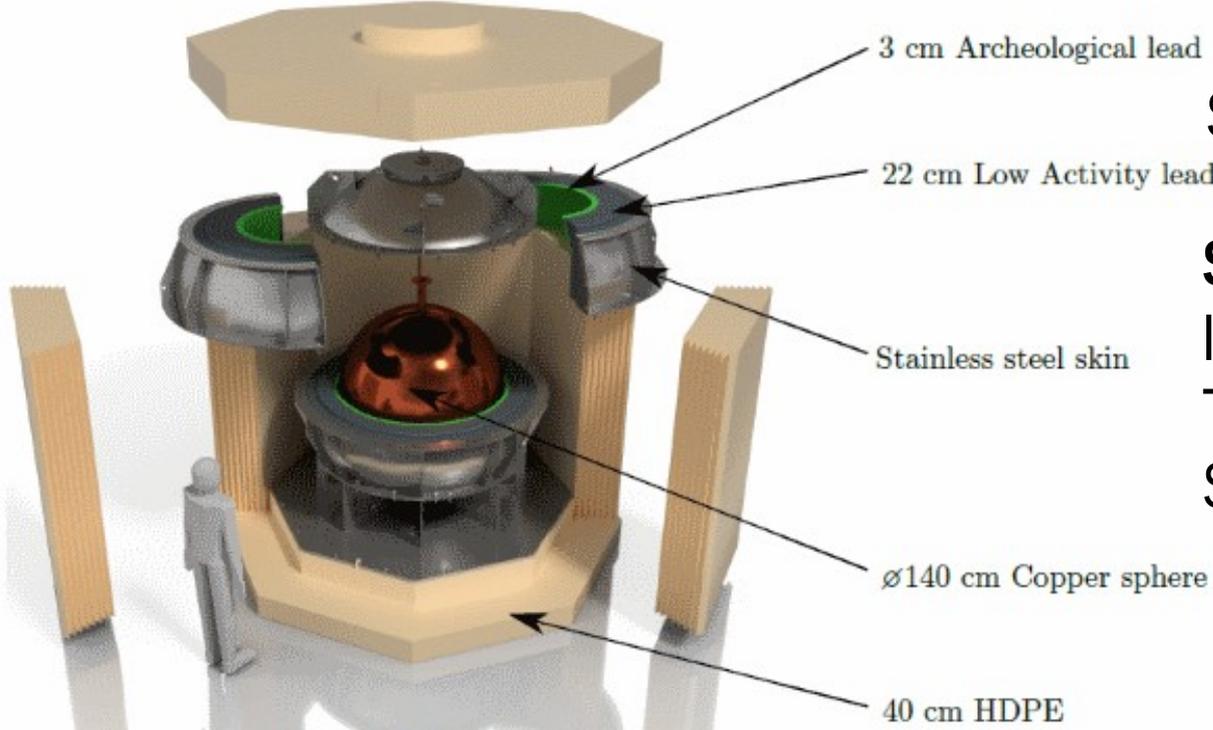
## 2017-2019: 3 bar He/CH<sub>4</sub>

# Sensitivity curve



# From 2020: NEWS-G @ SNO

<https://news-g.org/news-snolab/>



## Goal:

Gas H, He with  $P \leq 10$  bar  
Sensitivity to cross section of  $10^{-41}$  cm<sup>2</sup>

## Status:

In construction  
Technical design ongoing  
Space at SNO assigned



## IN2P3 contributions:

Same as ANR +  
Signal processing (since June 2018)

# Quenching factor measurements

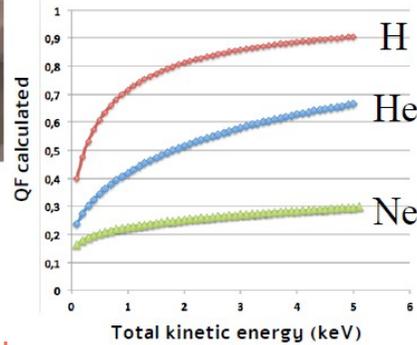
Goal: measure QF down to 500 eV ion energy using the Grenoble MIMAC facility for H, He, Ne, CF<sub>4</sub>, Ar, Xe at various pressures

Saclay, Grenoble, Thessaloniki, Queen's-Kingston

rg19, 1 bar He+5%CH<sub>4</sub>, HV=1971 V

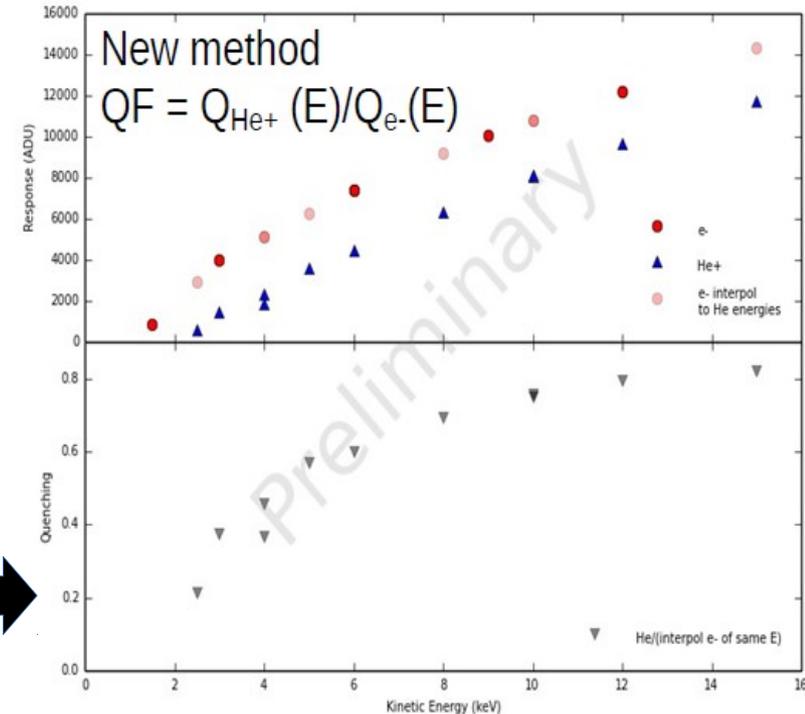


QF depends on  
- ion type  
- gas mix  
- gas pressure  
=> need 1 dedicated  
measure per detection  
setup



Range of interest for Dark Matter

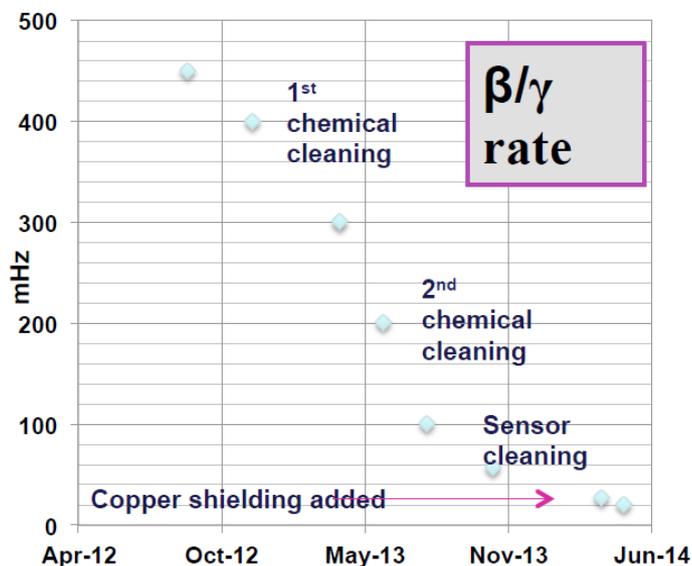
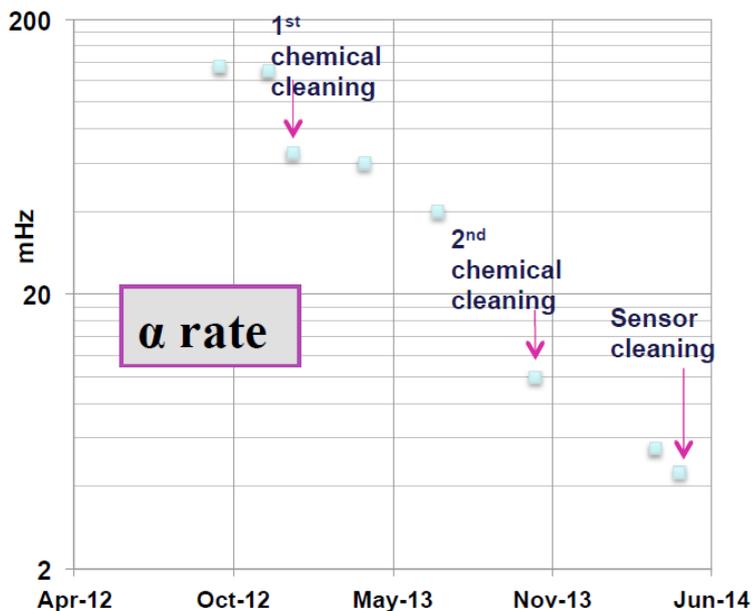
Earlier measurements D. Santos et al., ArXiv:0810.1137:  
Differences between measurements and calculations



# Improvement of Cu induced background

For Radon cleaning inside gas => huge expertise of the global community

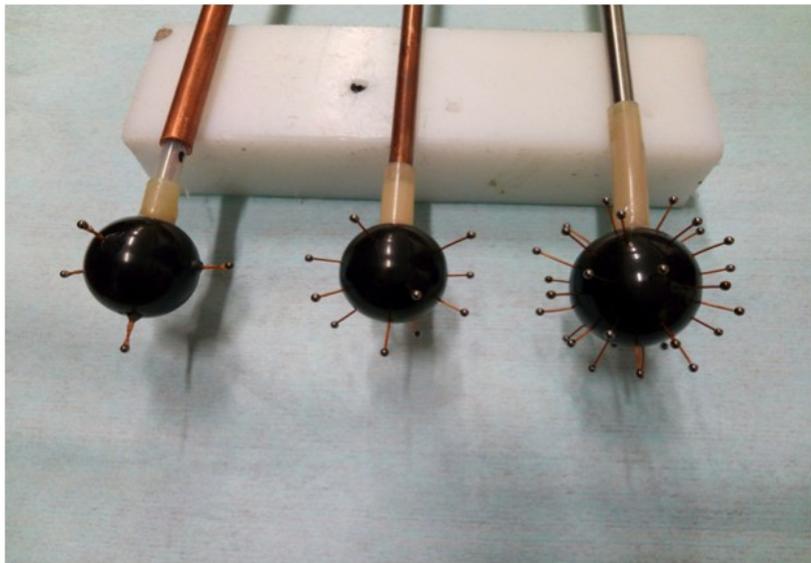
Goal: to remove  $^{210}\text{Po}$  - $^{210}\text{Pb}$  from Cu => chemical cleaning => - 10  $\mu\text{m}$



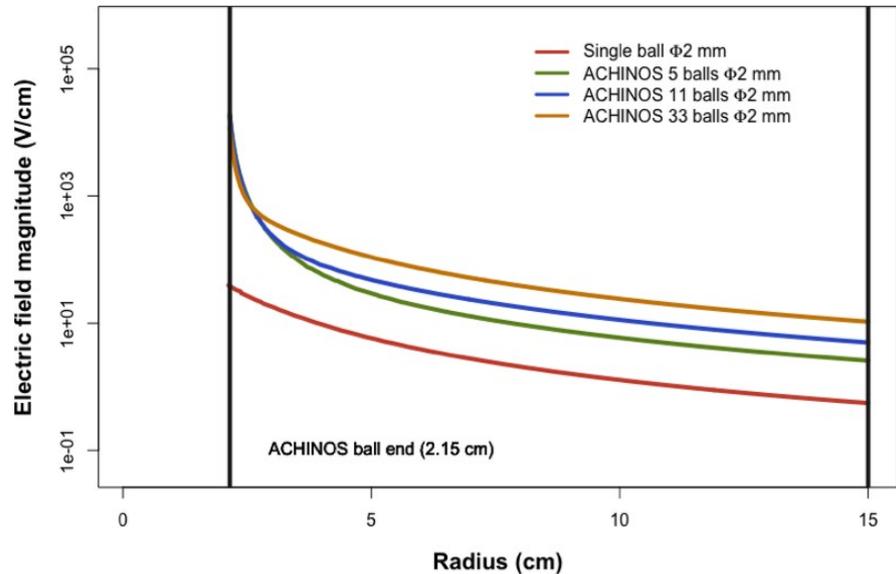
I. Giomataris Journée Matière Sombre Paris 2016



2017 => New method: electrolysis of Cu => - 10  $\mu\text{m}$  + 500  $\mu\text{m}$  (possibly up to +1 cm ...)  
=> Radio-purity of electro-deposition at the level of 10 nBq/kg

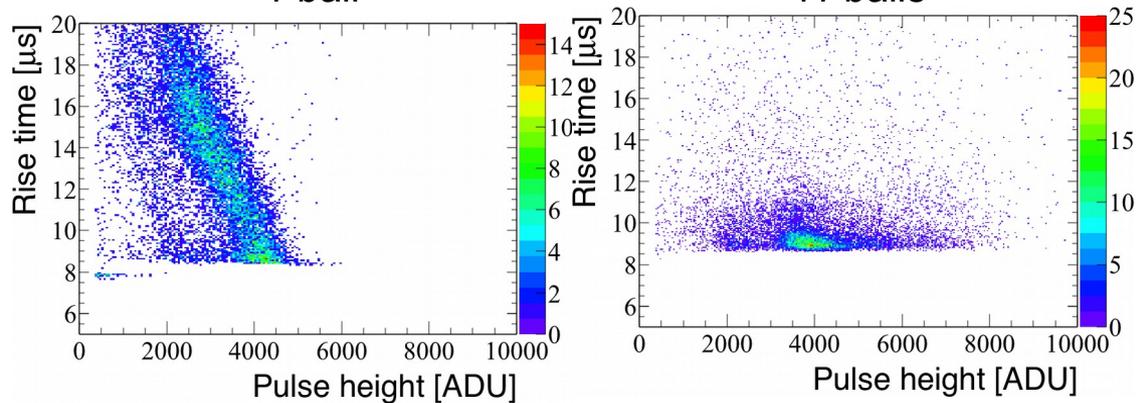


A. Giganon arXiv:1707.09254



Improvement of drifting  
 $\Rightarrow$  Multi-anode "Achinos"  
 $\Rightarrow E(r) \nearrow \Rightarrow$  Drift Time  $\searrow$

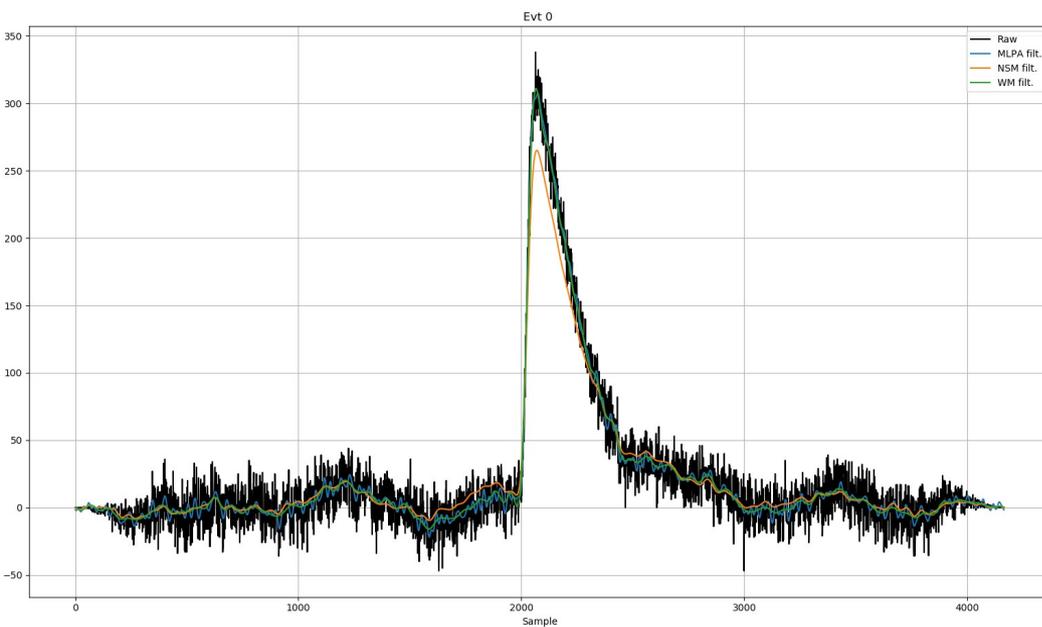
He:Ar:CH<sub>4</sub> (80:11:9) - 640 mbar - <sup>55</sup>Fe - 2-mm anode ball



+ Track recognition (spatial coverage of each individual anode ball)

# Improvement of S/N through signal processing

SUBATECH since June 2018

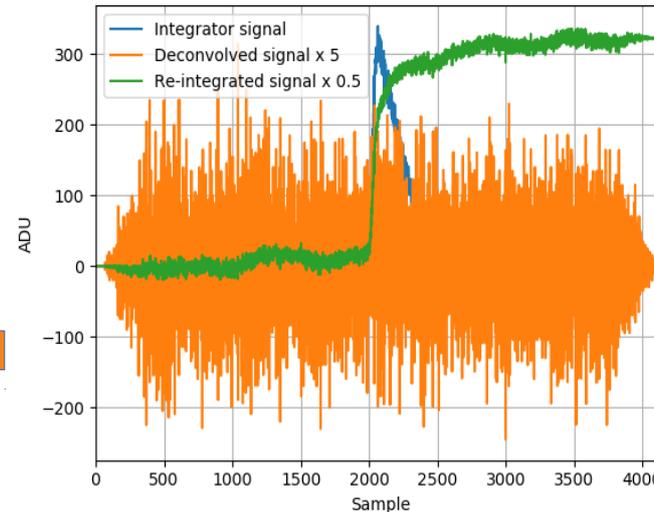


Filtering

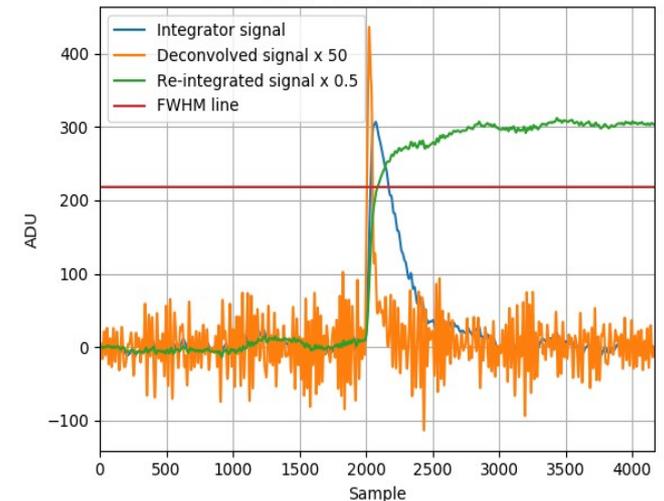
+ Regularization

+ Deconvolution

Evt 0 - No filtering - RC=140 micro s



Evt 0 - RC=140 micro s



- Reveals the distribution of charges
- Corrects the ballistic deficit

# R&D R2D2

## (Rare Decays with Radial Detector)

### SPC development for $2\beta\beta 0\nu$ search

J. Busto, C. Cerna, F. Druillole, A. Dastgheibi-Fard, C. Jollet, I. Katsioulas, I. Giomataris, M. Gros, P. Lautridou, A. Meregaglia\*, X. F. Navick, F. Perrot, F. Piquemal, A. Rebi, M. Roche, B. Thomas, M. Zampaolo

*CPPM, CENBG, LSM, IRFU, SUBATECH*

*Master-project initiated in 2017*

**2 objectives: to improve energy resolution & radioactive background compared to last experiments**

# Use of $^{136}\text{Xe}$ at high pressure

## Avantages

- High density is desirable to contain event
- Moderate density helps to reject Compton backgrounds

EXO-200 ~3-4 % FWHM

NEXT: 1% FWHM (20 bar + light detection)

KamLAND-Zen: 8% FWHM

XENON ~ 3% FWHM

*A. Bolotnikov, B. Ramsey / Nucl. Instr. and Meth. in Phys. Res. A 396 (1997) 360-370*

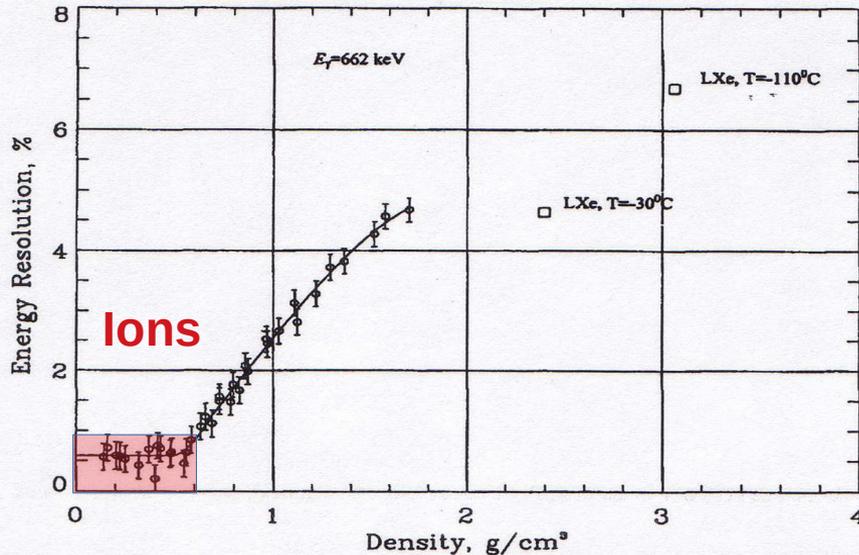


Fig. 5. Density dependencies of the intrinsic energy resolution (%FWHM) measured for 662 keV gamma-rays.

**R2D2 objective**

**Reach < 1% FWHM @ 40 Bar  
(Impact on ROI width)**

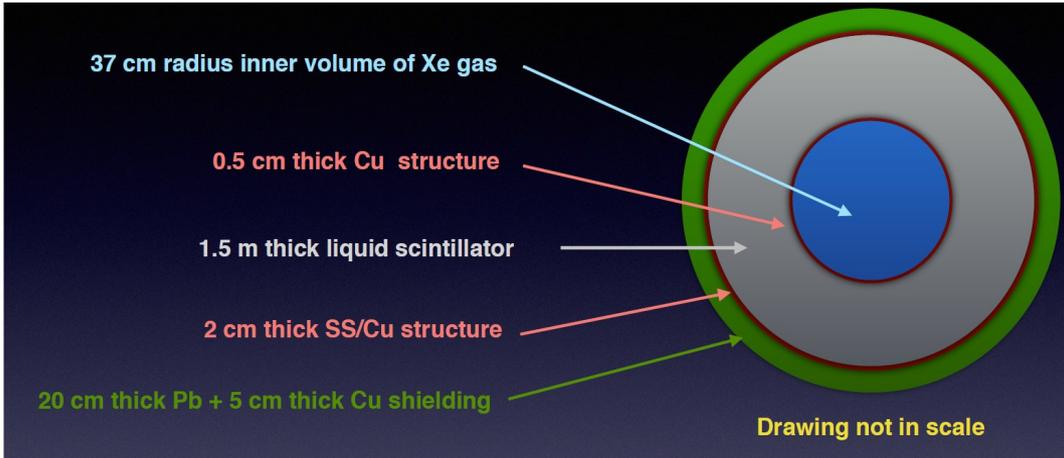
**Issue:**  $\Delta E/E$  after avalanche ?

Other gas (MoF6 ...) also envisaged in the future

# Simulations of Background

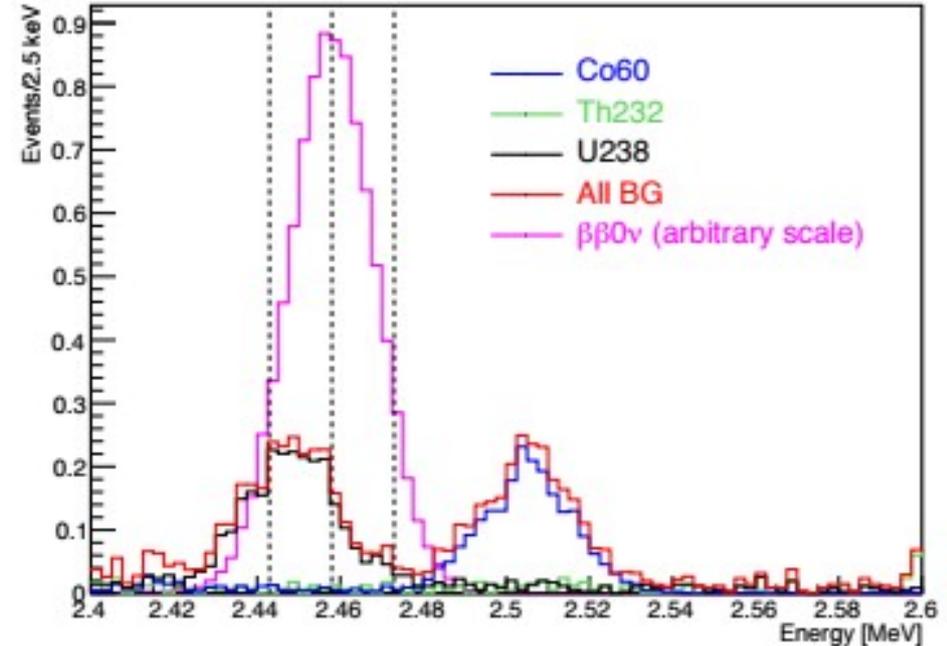
## Settings

- $\Delta E/E = 1\%$  FWHM @  $Q_{\beta\beta}$  of 2.458 MeV
- 50 kg  $^{136}\text{Xe}$  @ 40 bar ( $\phi = 74$  cm)
- 10  $\mu\text{Bq/kg}$  Cu activity for all sources of background ( $< 2 \mu\text{Bq/kg}$  available)



$\Rightarrow$  In ROI: 2 background evt/y  $\Rightarrow T_{1/2}^{0\nu}$   
 $= 2.5 \cdot 10^{25}$  Y,  $\langle m_{\beta\beta} \rangle < (160 - 330)$  meV

A. Meregaglia JINST 2018



ROI width:  $Q_{\beta\beta} \pm 0.6\%$   
 1 year data tacking

# Roadmap

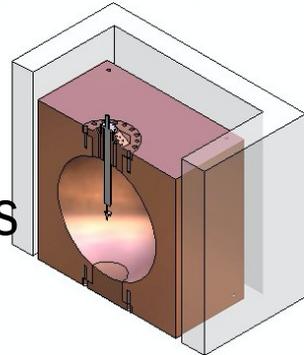
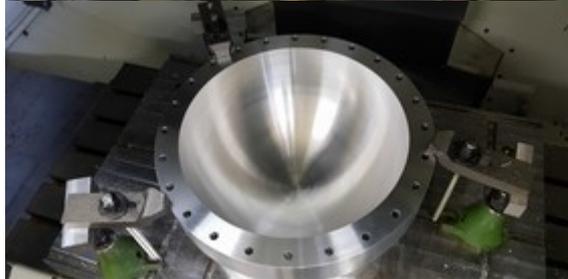
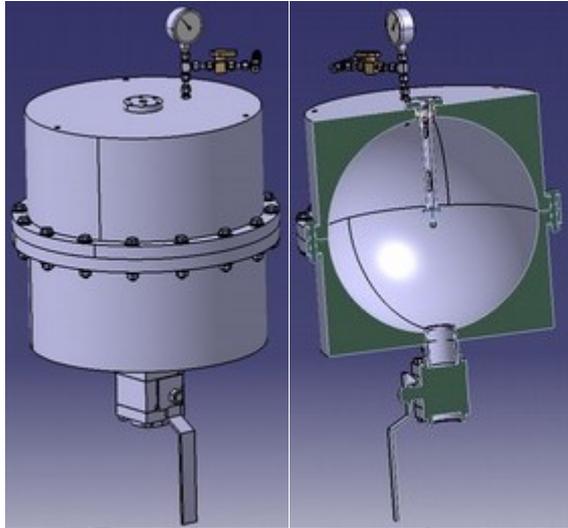
**2019:  $\Delta E/E < 1\%$  FWHM @ CENBG**

(SPC construction ( $\phi = 40$  cm) + Gas recovery system + Anode optimization for 6000 V- HV + Long track & 2-blob recognition,...)

- **With ArP2 @ 1-40 bar**
- **With natural Xe @ 1-40 bar** (+ Test of Rn contamination)
- **Search for fundings (ANR, ERC...) in parallel**

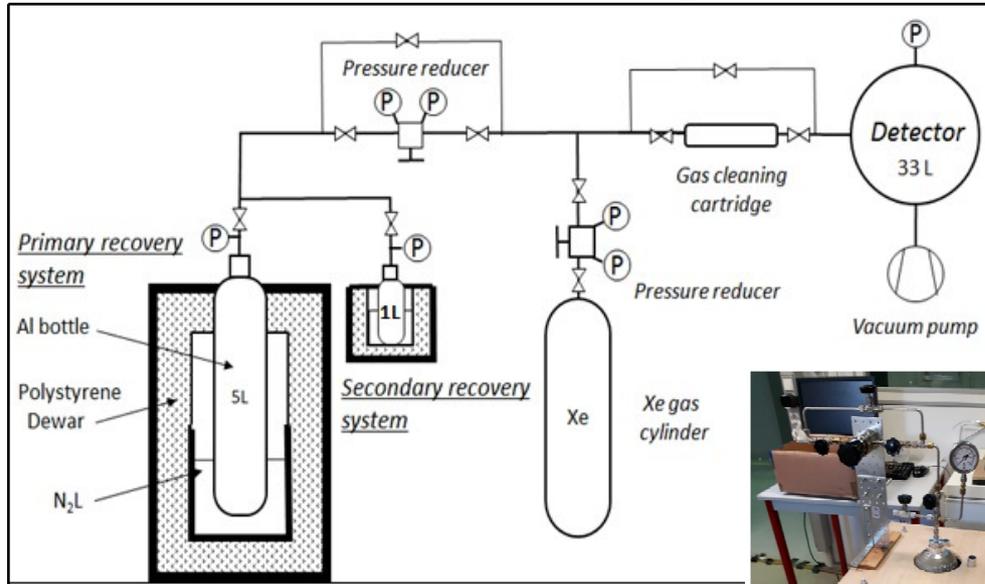
**2020: “zero- background” @ LSM with SPC**  
( $\phi = 60$  cm, 27 kg  $^{136}\text{Xe}$ ) => first results of physics

=> Prospect: A ton-scale detector (possibly modular) with zero-background



# Recovery & Cleaning system

## Recovery by cryopumping @ N<sub>2</sub>L temperature



- Very clean technique (no mechanical pump)
- Simple and low cost

### 2 recovery tanks

- Primary => 5L Al tank -> 660 L Xe gas  
(15% of Xe remains in the recovery tank)
- Secondary => 1L Al tank  
(3% of Xe remains in the recovery tank)



## Cleaning



Hot getter



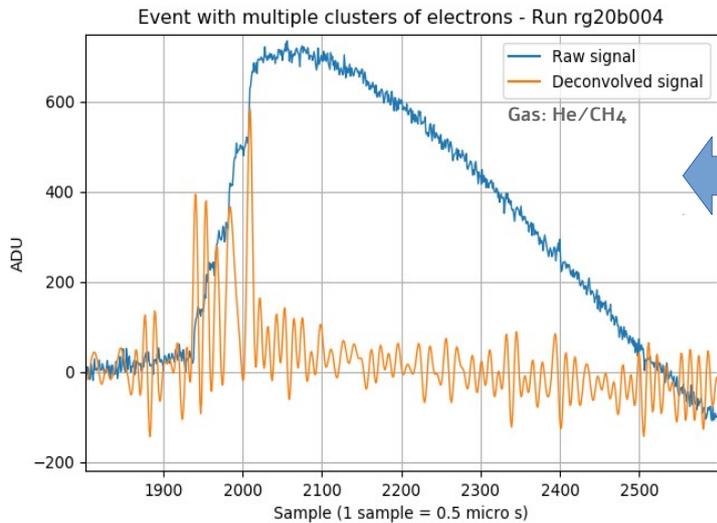
Ambient absorber

- Very high gas purity is required (O<sub>2</sub> and H<sub>2</sub>O)
- High purity Xe (Air Liquide, Linde) : 99.999 %  
=> 1 ppm O<sub>2</sub>

### Adsorbent ambient cartridge :

- Supelco® O (< 0.5 ppm O<sub>2</sub>)
- GateKeeper® GPU (< 1 ppb O<sub>2</sub>)
- SAES® (< 0.1 ppt O<sub>2</sub>)

Hot getter : SAES® (< 1 ppb O<sub>2</sub>)

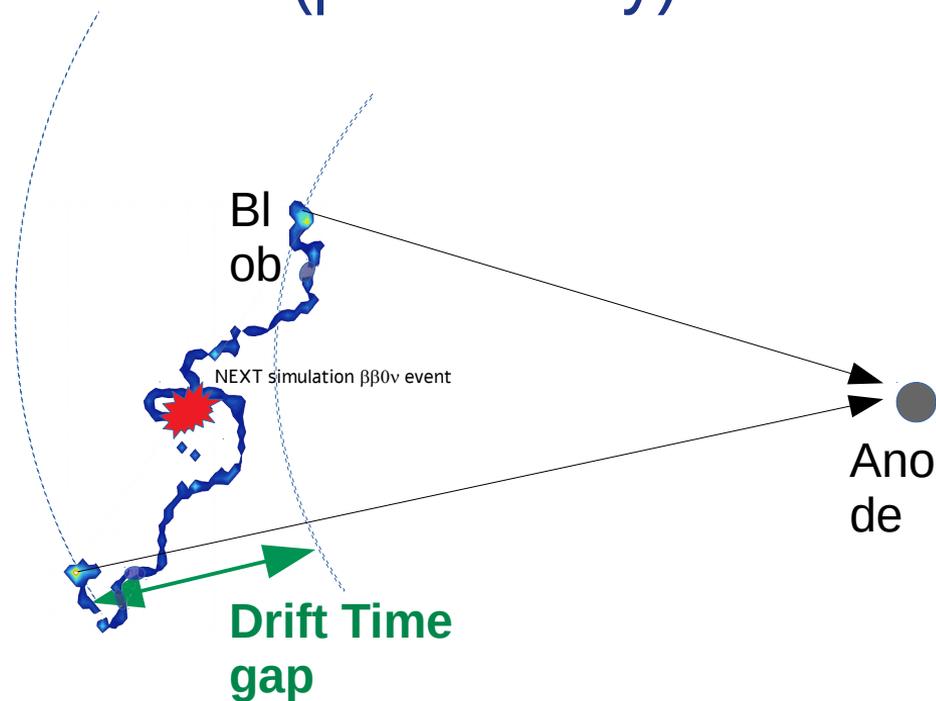
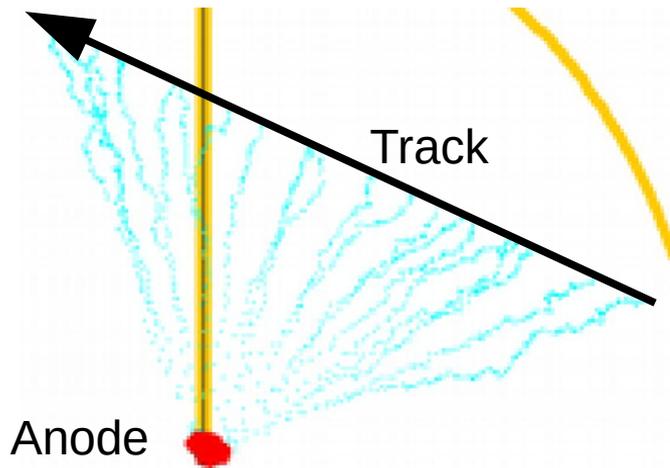


Experimental observations

& After signal processing

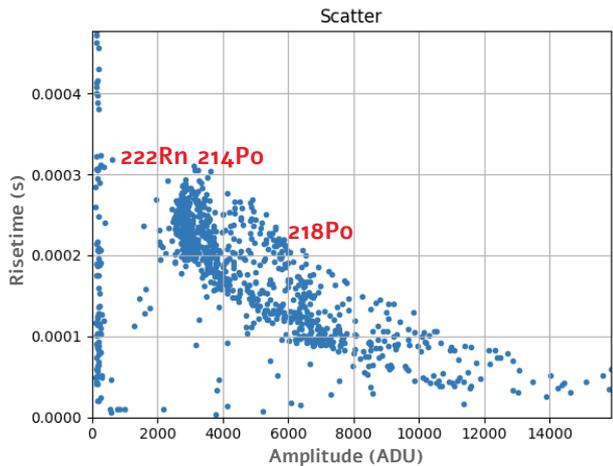
Interpretation: Long track signal ?

P. Gros UCLA Dark Matter conference 2018



New possibility ?  
 Ability to recognize  
 multi-blob events ?  
 (preliminary)

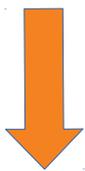
He/CH4 1 Bar + Rn source



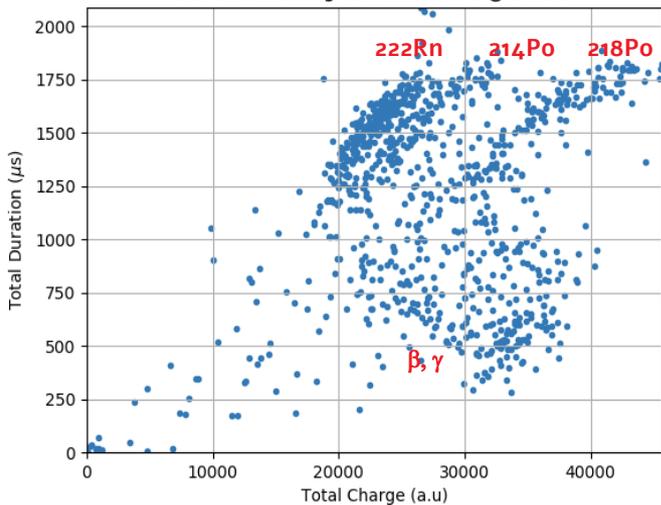
# New possibility ?

## Ability to recognize the nature of interactions ? (preliminary)

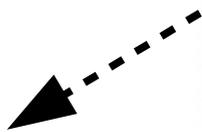
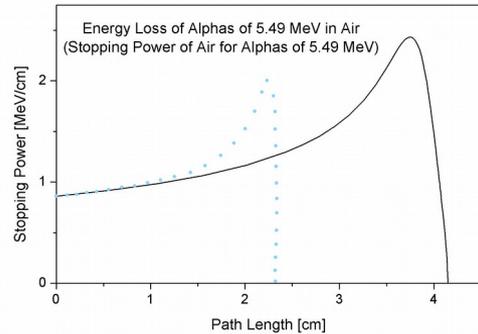
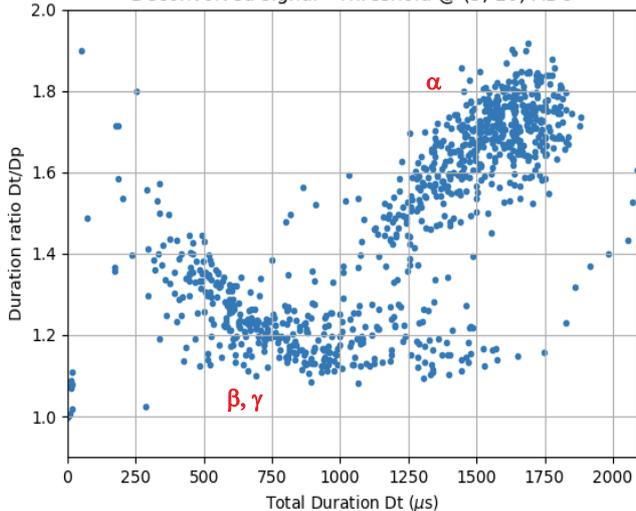
After signal processing  
+ PSD method



Deconvolved signal - Threshold @ 5 ADU



Deconvolved signal - Threshold @ (5, 10) ADU



Signal content linked to the  
Bragg curve ?

# Organization

