

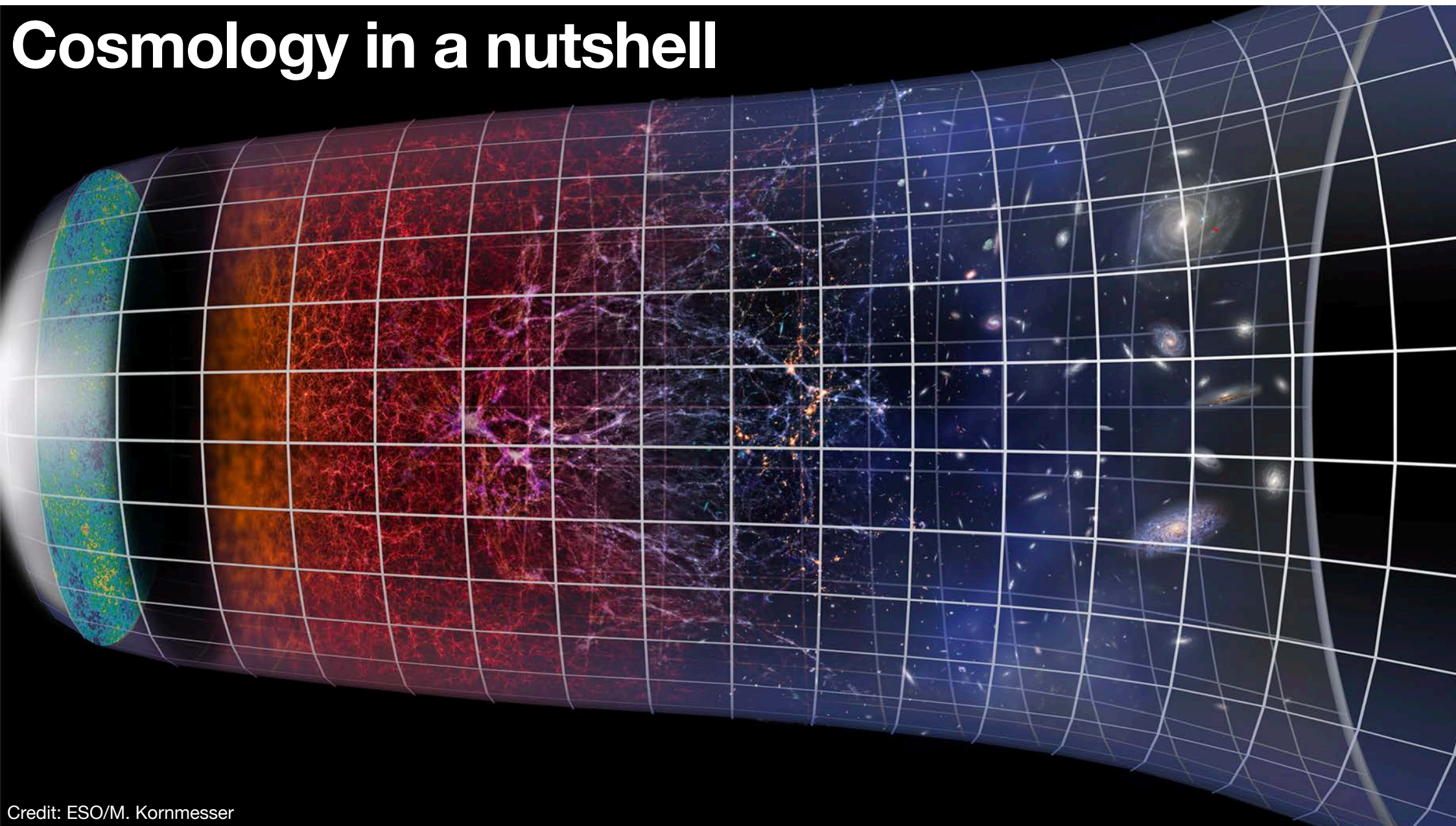


# Theoretical Cosmology

Conseil Scientifique de l'IN2P3

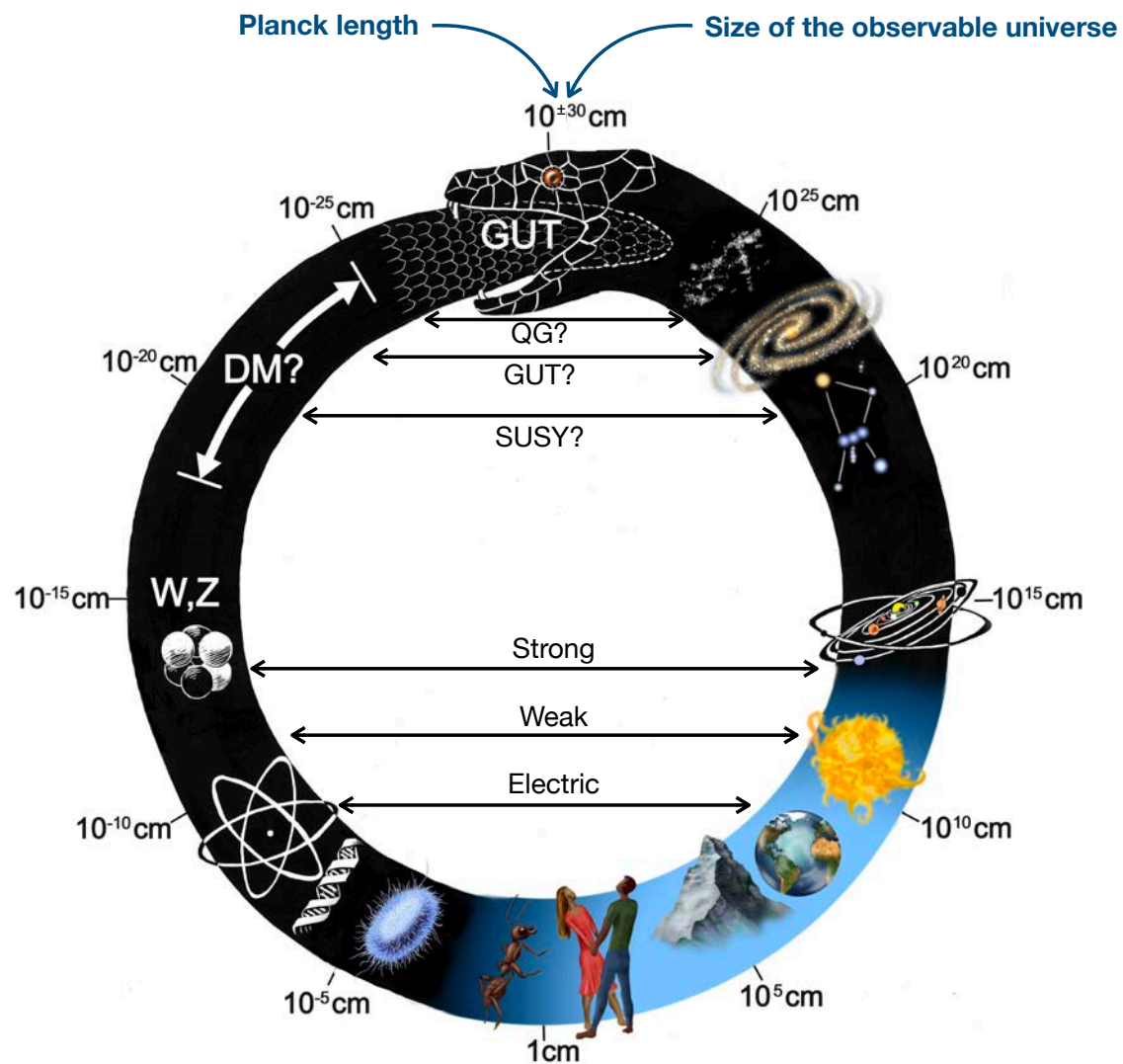
Vincent Vennin, 29th June 2021

# Cosmology in a nutshell



Credit: ESO/M. Kornmesser

# Cosmology in a nutshell



*The cosmic Uroboros*

# Summary of the workshop:

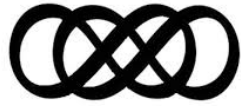
*Physique Théorique des*  .

A SELECTION OF TOPICS IN THEORETICAL COSMOLOGY

Program and slides available at: <https://indico.in2p3.fr/event/23540/overview>

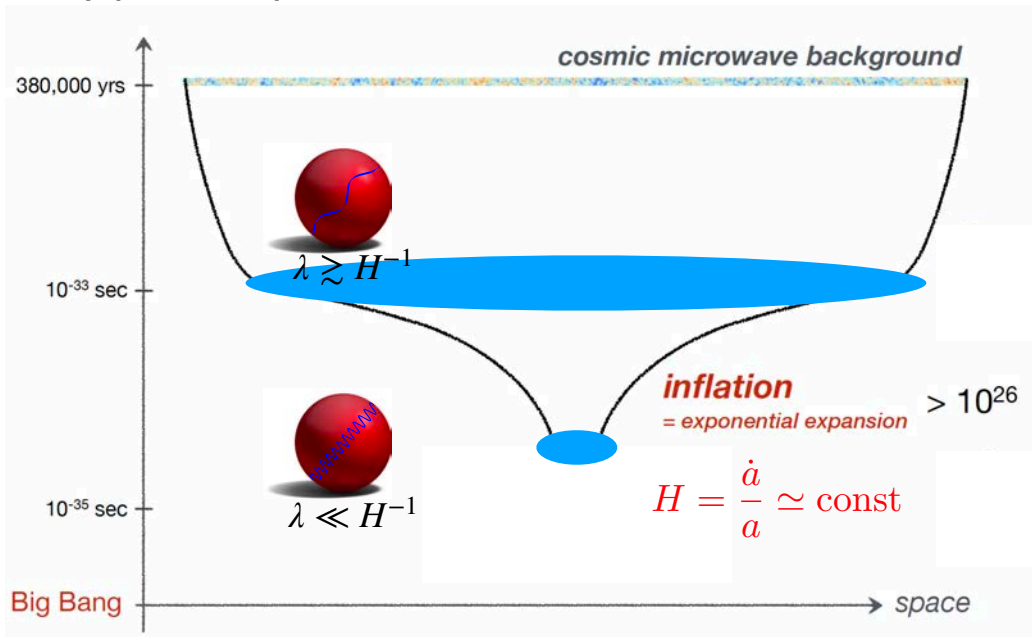
**See written report for further details about specific contributions from IN2P3 researchers**

# Early Universe



## ***Inflation: a giant microscope***

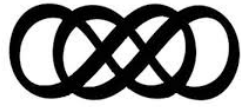
a tiny patch of space becomes the entire observable universe



### Open issues:

- What is the inflaton?
- To which extension of the standard model does it belong?
- At which energy did inflation occur?
- How did the inflaton transfer its energy to the standard model particles?
- Was it the only field around?
- Did super-Planckian degrees of freedom play a role?
- What does inflation tell us about the nature of Quantum Mechanics, and about the interplay between Quantum and Gravitational physics?
- What about alternatives to inflation?
- How to best constrain inflation from the CMB / LSS / GWs / etc?

c.f. presentation by Sébastien Renaux-Petel (Institut d'Astrophysique de Paris): "Early Universe Cosmology"



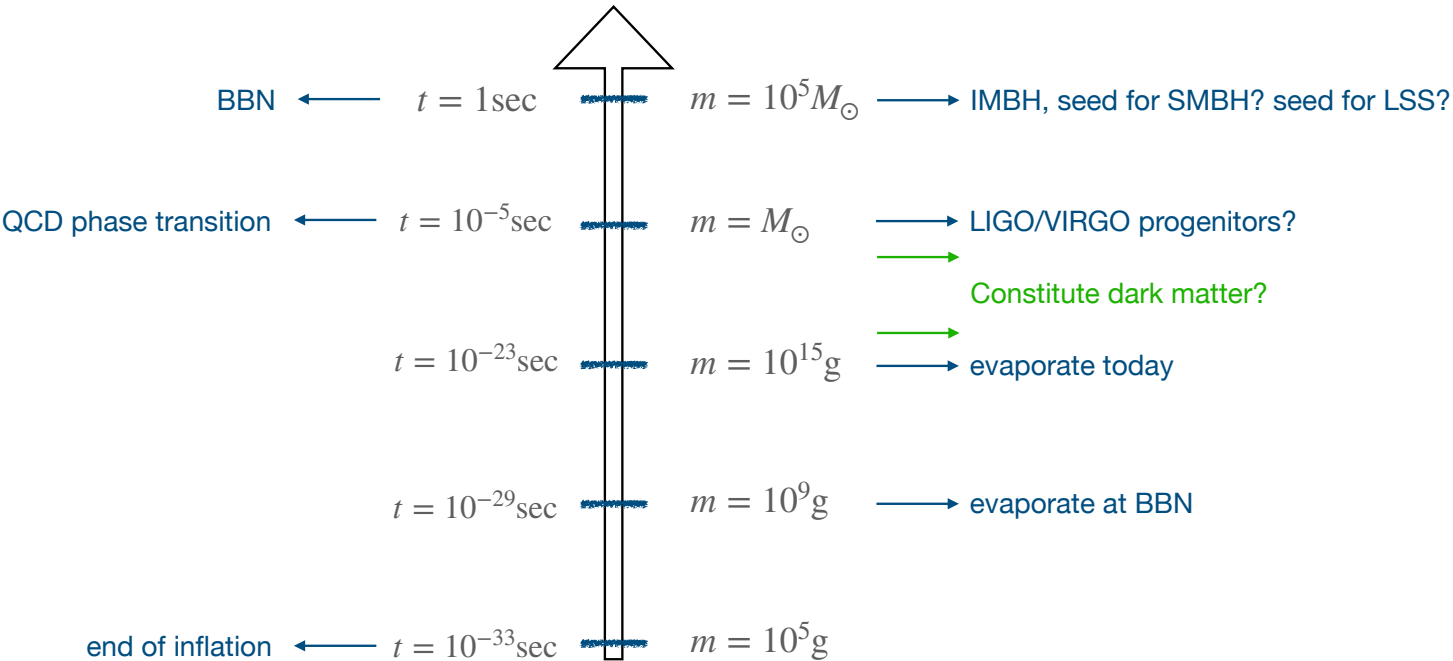
# Primordial Black Holes



## Possible origins:

- Collapse of large primordial over densities
- Phase transitions
- Collapse of topological defects

$$m_{\text{PBH}} \sim m_{\text{Pl}} \frac{t}{t_{\text{Pl}}} \sim 10^{38} \text{g} \frac{t}{\text{sec}}$$

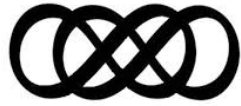


## Open issues:

- Formation mechanisms?
- How to compute their abundance?
- Role of Hawking evaporation?
- Testing GR?
- ...

c.f. presentation by Alexandre Arbey (Institut de Physique des 2 infinis de Lyon): "Primordial Black Holes"

# Modified gravity



- So far, **GR** seems compatible with all observations.
- Several motivations for exploring **modified gravity**
  - Quantum gravity effects
  - Understand cosmological acceleration (or possibly dark matter)
  - Explore alternative gravitational theories
  - **Testing gravity**
- Modified gravity actively studied in two main contexts:
  - In cosmology (alternative to the cosmological constant, exotic early Universe models)
  - In astrophysics: compact objects (black holes, neutron stars)

It is rather difficult to modify gravity:

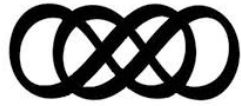
1. The theory must be **internally consistent** (e.g. no problematic instabilities)
2. The theory **must look like GR** in all regimes where GR has been tested
  - Lab tests, Solar system, Binary pulsars (and now binary BH)
3. Hopefully (but not necessarily), the theory should **account for the observed acceleration** and exhibit some **distinctive signatures**.

## Possible signatures

- Speed of GWs different from  $c$
- Evolution of cosmological perturbations different from GR
- GWs from binary merger with different waveforms (quasi normal modes)
- Different internal structures for compact objects (strong-field regime)

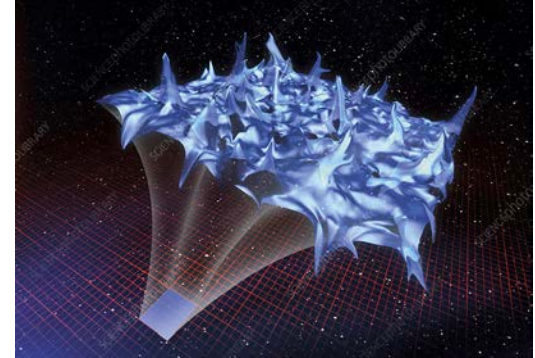
**c.f. presentation by David Langlois (Laboratoire Astroparticules et Cosmologie): Modified Gravity in Cosmology and Astrophysics**

# Quantum gravity



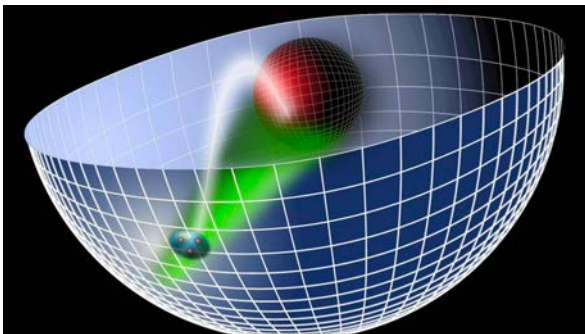
## Problem:

- We do not know how to combine gravitational and quantum physics
- There are a number of places in cosmology where both play a role
- Need to modify GR? QM? Both?



## Various approaches:

- String theories
- Loop quantum gravity
- Approaches inspired by the holographic principle, AdS/CFT, emergent gravity etc
- ...



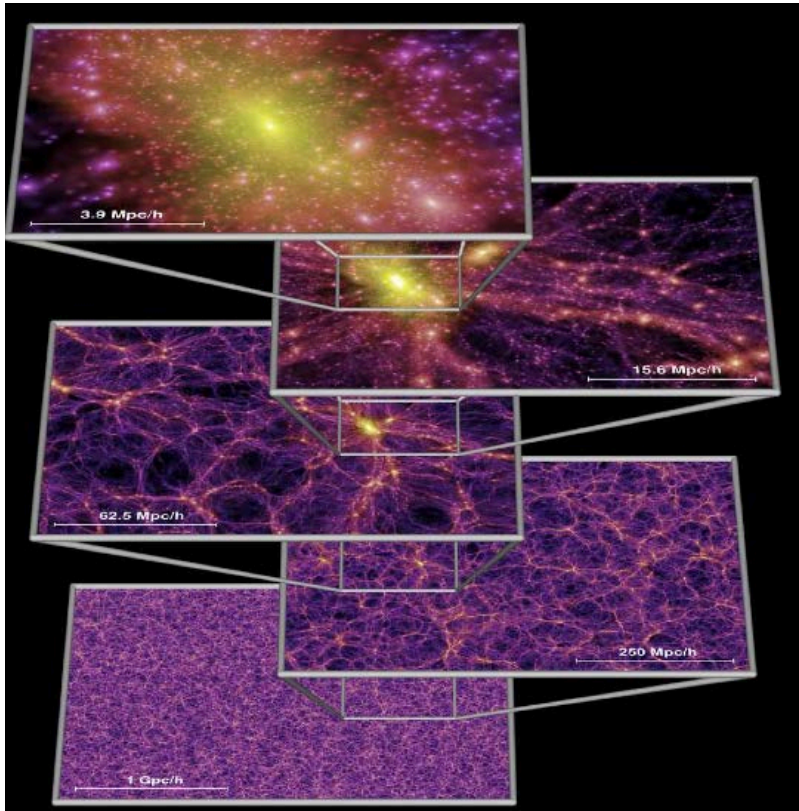
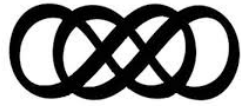
## Possible applications: a better understanding of

- Inflation
- Dark energy
- Black holes
- ...

c.f. presentation by Elias Kiritsis (Laboratoire Astroparticules et Cosmologie): Emergent Gravity and Cosmology



# Cosmological Structure Formation



To account for the organisation of matter (visible and dark) on the largest scales (roughly: galactic scales upwards 0.1 - 10 Mpc)

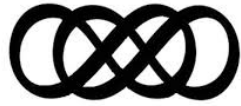
## Goals:

- Account for the formation of structures in our Universe
- Constrain cosmology (cosmological parameters, theory of gravity, etc)
- Reconstruct initial conditions (primordial fluctuations) and thus constrain early Universe

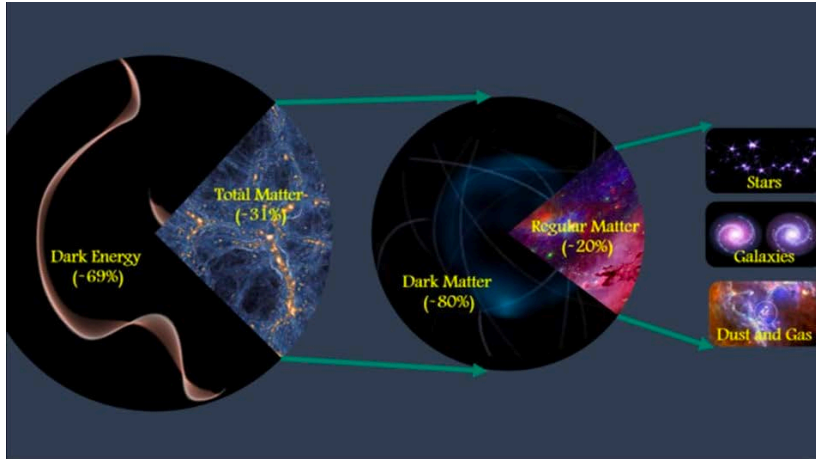
## Technical challenges:

- Role of baryonic & AGN feedback
- Role of non-linearities
- Heavy numerical methods (N-body codes tested with scale-free models, large volumes, MCMC with many realisations)
- Statistical characterisation of the cosmic web (clustering, intrinsic alignment, etc): what should be measured?

c.f. presentation by Michael Joyce (Sorbonne Université): Non-linear cosmological structure formation



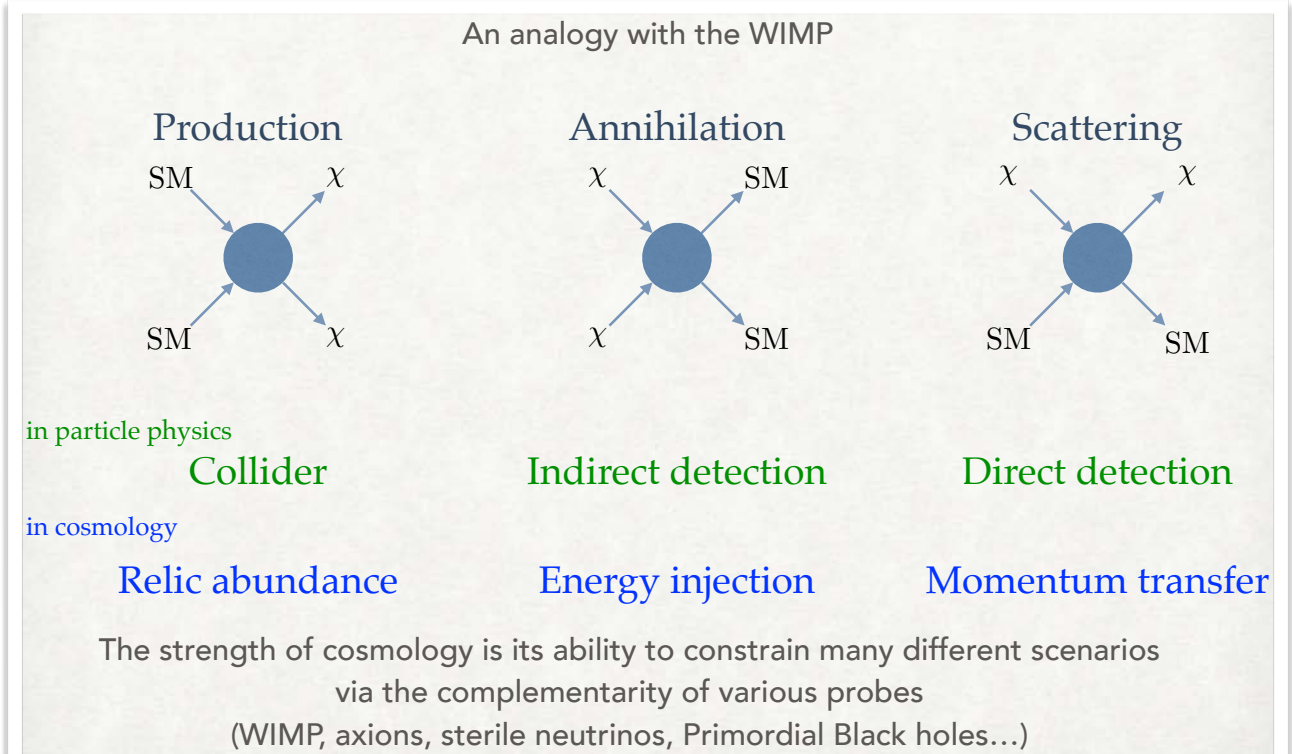
# Dark matter



Many gravitational clues for the existence of Dark Matter on a variety of scales.  
Description “CDM” is purely parametric: can we probe the nature of DM?

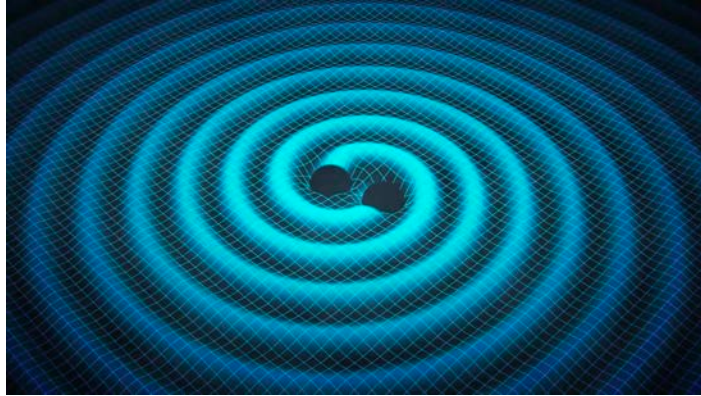
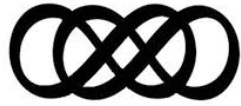
## Technical challenges:

- extract “cosmological” & “particle physics” information from large surveys with large astrophysical uncertainties.
- Broad phenomenology: complementarity of probes over a wide-variety of scales and times.
- accurate modeling of small-scale physics + statistical analysis.  $N$ -body? EFT of LSS?
- Do  $\sigma_8$  and  $H_0$  tensions point towards interactions/decays in the dark sector?



c.f. presentation by Vivian Poulin (Laboratoire Univers et Particules, Montpellier): Dark matter in cosmology

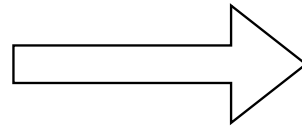
# Gravitational Waves Cosmology



Goal: Use GW observations to probe the cosmic expansion and test different cosmological scenarios / models

Standard Sirens: GW events that can be used as absolute cosmological distance indicators

- Luminosity distance estimated from GW signal
- Redshift obtained from EM observations or features in the mass distribution of GW sources



Fit the distance-redshift relation

- Constrain cosmological parameters
- Test GR on cosmological scales
- Probe inhomogeneities?

Example of multi-messenger astronomy

**c.f. presentation by Nicola Tamanini (Laboratoire des deux infinis Toulouse): Gravitational-waves cosmology, a new arena to test the dark universe**

# Specific aspects of the field

## Scientific coverage:

- Spans a large range of scales —> multidisciplinary, interface-driven
- Strong interactions with all other components of IN2P3: formal theoretical physics, particle physics, nuclear physics, astroparticles etc

## Methodological coverage:

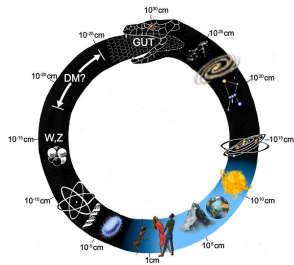
- Wide spectrum of methods: algebraic, analytical, topological, geometrical, QFTs, EFTs, numerical, statistical, etc...
- Fosters mobility of researchers across topics
- Stimulates research at the interface

## Institutional coverage:

- Various **IN2P3-affiliated laboratories**
- Laboratories with **other main affiliations** (INP, INSU)
- Hiring from different CNRS sections: 01 (IN2P3), 02 (INP), 17 (INSU); besides university positions
- Substantial amount of collaborators based abroad, which results into various international partnerships. Research without borders!

- **APC Paris** (Chiara Caprini, Nathalie Deruelle, Eric Huguet, Elias Kiritsis, David Langlois, Jihad Mourad, Francesco Nitti, Jacques Renaud, Julien Serreau, Danele Steer, Vincent Vennin, Cristina Volpe)
- **IJCLab Orsay** (Eugeny Babichev, Christos Charmousis, Yann Mambrini, Karim Noui, Bartjan Van Tent)
- **IP2I Lyon** (Hubert Hansen, Jerome Margueron, Alexandre Arbey)
- **IPHC Strasbourg** (Michel Raush de Trautenberg)
- **L2IT Toulouse** (Nicola Tamanini)
- **LAAP Annecy** (Tania Regimbau)
- **LPNHE Paris** (Michael Joyce)
- **LPSC Grenoble** (Aurelien Barrau, Killian Martineau)
- **LUPM Montpellier** (Karsten Jedamzik, Julien Larena, Julien Lavalle, Vivian Poulin)

- **INP**: LAPTH Annecy, L2C Montpellier, CPT Marseilles, LPTHE Paris, LPENS Paris
- **INSU**: IAP Paris, LUTH Meudon, IAS Orsay



# Specific aspects of the field

- Large amount of topics being covered in spite of moderate number of involved scientists
- Various connection with experimental projects
- Strong international visibility
- Pronounced ability to embrace new topics, and to show versatility in a quickly-developing field
- High potential for breakthrough, wide impact across communities (3 out of the last 4 Nobel prizes were given to Cosmology!)



- Interaction rate within the French community is sometimes limited (too few recruitments to cross the critical-mass threshold in some topics given growing activity and broad scope of the field; lack of theory-specific national platform)
- Performance with funding applications sometimes suffers from schemes being not adapted to theory, and theory being dispatched across several items