



# KM3NeT/ORCA

Conseil Scientifique de l'IN2P3 26/10/2021

Mathieu PERRIN-TERRIN on behalf of the KM3NeT groups of APC, CPPM, IPHC, LPC, Subatech, LUPM



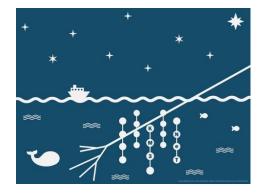






## Outline

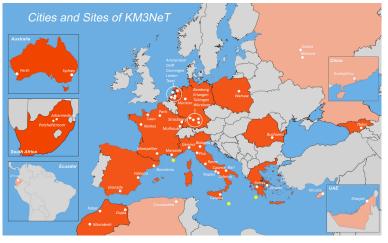
- Introduction
  - KM3NeT Collaboration
  - KM3NeT Technology
  - French implication
- Status of the ORCA construction
  - Detectors construction
  - Deployments (past and future)
- Status of the Physics Analyses
  - Analysis of the first data
  - Expected sensitivities
  - Tag Neutrino Beam from Protvino





#### Introduction

- KM3NeT Collaboration is building the next generation of v telescope
  - 56 institutes distributed across 17 countries
  - Successors of ANTARES which has been operational since 2007
- Two main scientific goals are addressed
  - determination of the v mass ordering (NMO)
  - discovery & observation of high energy neutrino source in the Universe
- Two detectors under construction
  - France: ORCA (Oscillation Research with Cosmics in the Abyss) optimise for GeV v
  - in Italy: ARCA (Astronomy Research with Cosmics in the Abyss) optimise for TeV v
  - Both detectors use the same technology
- Selected on the <u>ESFRI roadmap</u>



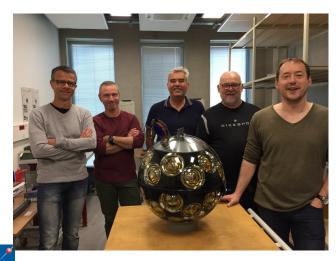
MONITORING OF ESFRI 2016 PROJECTS				
10	cientific Case	nplementation Case	Overall	
CTRIS	High	High	High	
ANUBIUS	ledium/High	Medium	Medium	
RIHS	Medium	Medium	Medium	
ST	High	Medium	Medium/High	
MPHASIS	ledium/High	Medium/High	Medium/High	
KM3NeT	Very High	High	ligh/Very High	



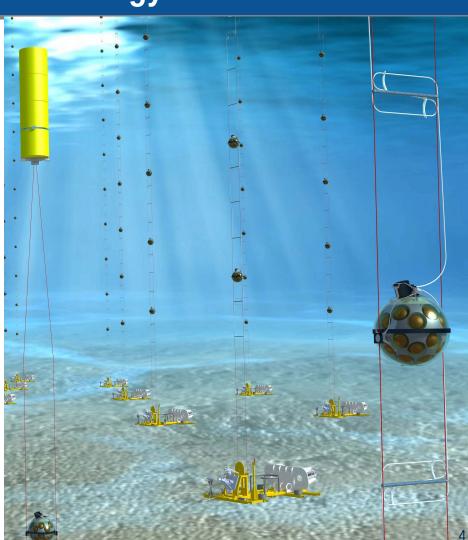
## **KM3NeT Technology**

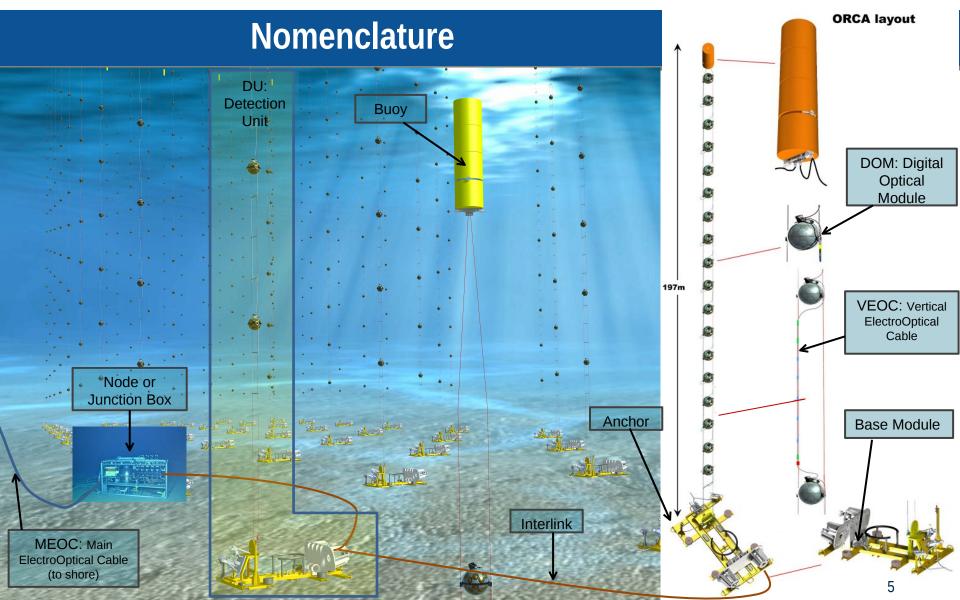
#### Array of Digital Optical Modules (DOM)

- Each DOM contains 31 PMTs (3")
- 18 DOMs form a detection line
- ORCA will consist of 115 lines
- ARCA will consist of 2x115 lines
- Spacing between DOM determines the detector energy threshold

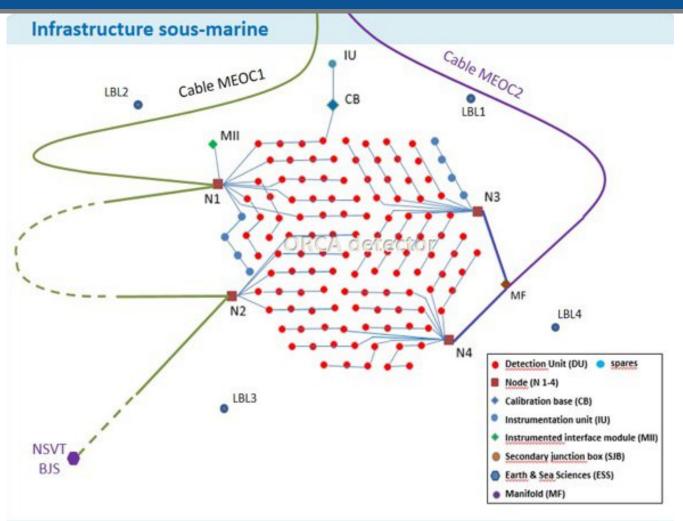








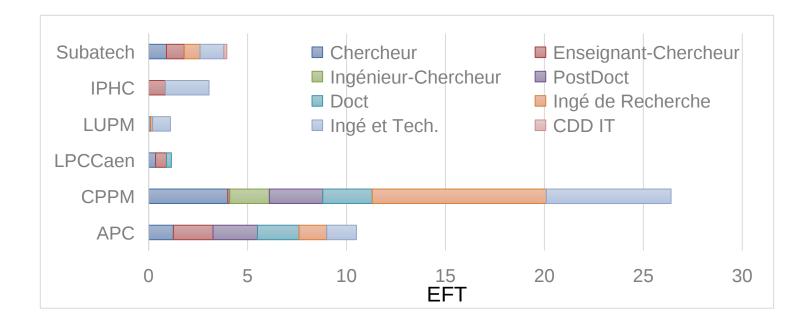
#### **ORCA Infrastructure**





#### **French Implications**

- Six IN2P3 Labs are members of KM3NeT
  - Two new members since 2017: LUPM and LPC-Caen





#### **French Responsibilities**

- Strong implication of the French groups
  - Cristal Collectif du CNRS for CPPM Team

Labo	Technical Responsabilites	Management Responsabilites
APC	Calibration Unit	KM3NeT Technical Manager ANTARES Spokeperson Chair Conf. Committee Co-coord Oscillation Physics
СРРМ	DU Integration Anchor Integration Sea Operations Control Room / Power System Instrumentation Unit Junction Boxes	KM3NeT Spokeperson Co-coord Oscillation Physics Co-coord Data Taking Co-coord Neutrino Astronomy Coord KM3NeT Outreach ORCA Site Manager
LPCCaen	DOM on VEOC Anchor integration	
LUPM	Base Module Integration	
IPHC	DOM Integration	
Subatech	DOM Integration	





#### **Technical Realization — DU Construction**

- Constructions since 2017
  - 76 DOMs [Subatech and IPHC]
  - 13 base modules [CPPM]
  - 16 DUs [CPPM]

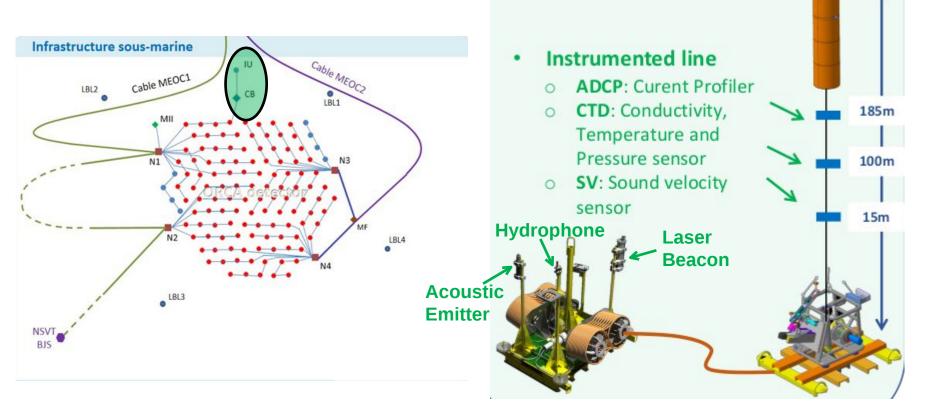






#### **Technical Realization — Calibration Instruments**

• Instrumentation Line & Calibration Base (CPPM & APC)

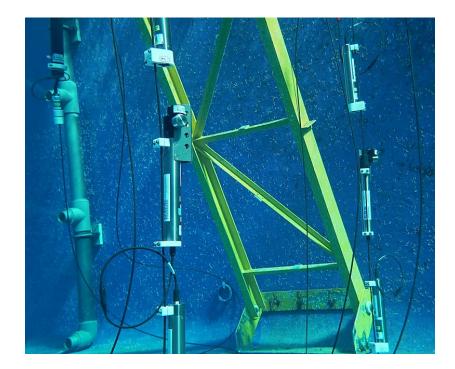




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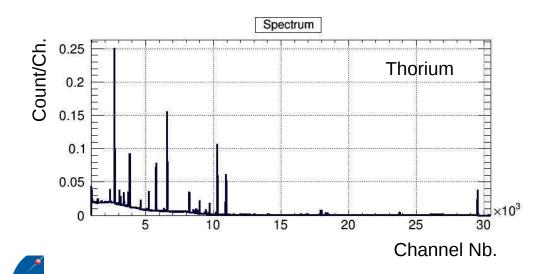






#### **Technical Realization — Earth and Sea Science**

- Gamma spectrometer (CPPM)
  - High purity Germanium (HPGe): excellent resolution
  - Cryogenic detector (70°K) marinized to operate at 2500m depth
  - Excellent γ energy resolutiom
  - Spectrum will be measured live !







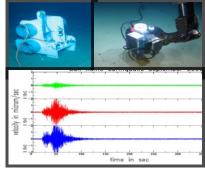
#### More Earth and Sea Science

Infrastructure is now a <u>platform</u>:

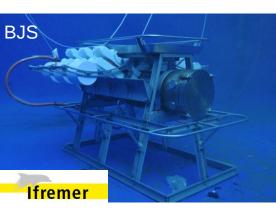
LABORATOIRE SOUS-MARIN PROVENCE MÉDITERRANNÉE



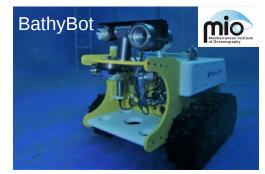
- Laboratoire Sous Marin Provence Méditerrannée



Seismometer







ctionnement de l'observatoire ences de l'environnement

·····)))

division technique

MII) câblé

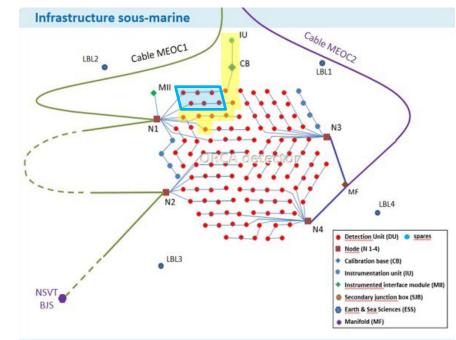
La communication sur la ligne entre le PC et les instruments se fait par transmission inductive

La communication entre le module et la ligne se fait par des modems en liaison acoustique

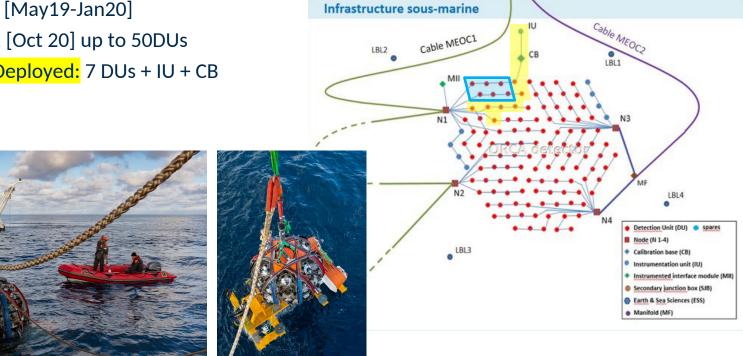
Ligne de mouillage ALBATROSS autonome

- Activities on the Infrastructure
  - MEOC1 & Node 1 repaired
  - 6 DUs [May19-Jan20]
  - Node 2 [Oct 20] up to 50DUs
  - To Be Deployed: 7 DUs + IU + CB





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@ PHOTOTHEQUE IN2P3 / CNI

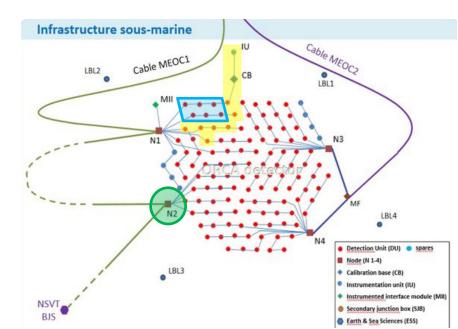


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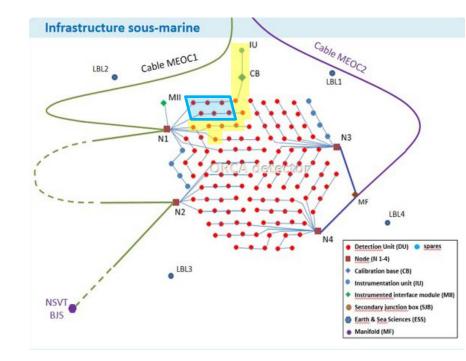






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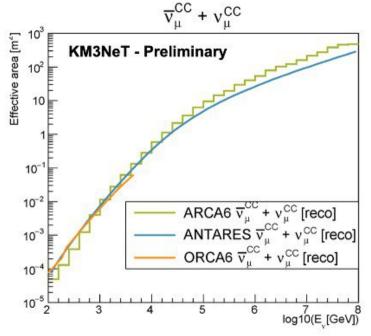




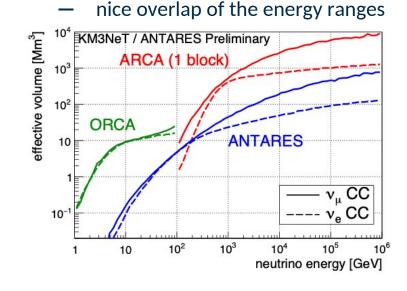


#### **ORCA6 ARCA6 and ANTARES**

- Current Status
  - 6 lines at ORCA
  - 6 (+2) lines at ARCA (+15 next April)

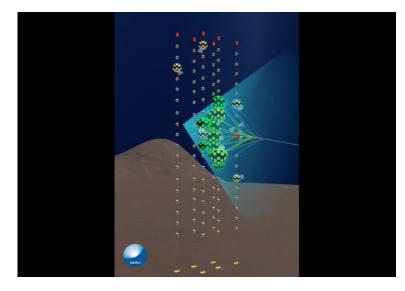


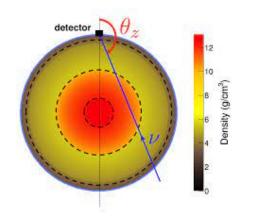
- Effective Area (i.e. efficiency)
  - larger than ANTARES on the whole energy range
  - ANTARES can be dismantled
- Full Detectors



#### **Neutrino Events**

• Event display of some up-going neutrinos

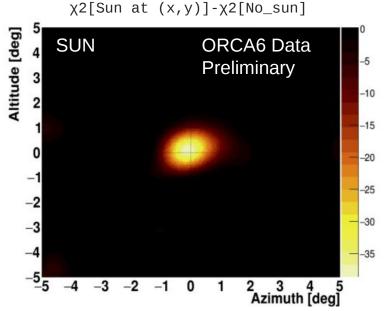






#### **Detector Absolute Pointing — Moon and Sun shadow**

- Sun and moon are screening the cosmic rays
  - depletion in cosmic ray muon rate at the object position
  - allows to probe the detector pointing
  - ORCA-6 already sees the sun (6.1  $\sigma$ ) and moon (4.0  $\sigma$ ) with an angular reso of 0.6°

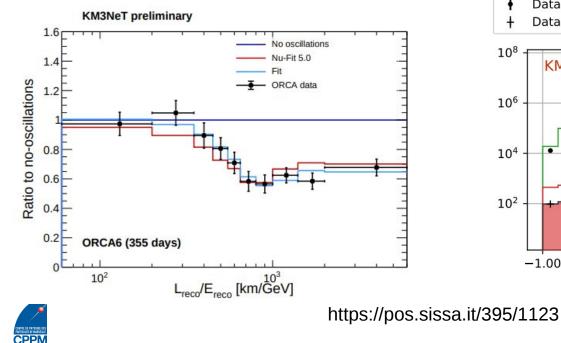


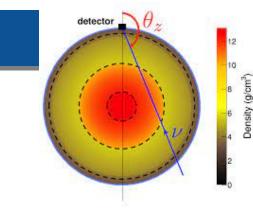
 $\chi^2$ [Moon at (x,y)]- $\chi^2$ [No\_moon] 5 Altitude [deg] MOON **ORCA6** Data 4 -5 Preliminary 3 -10-15 -20 -25 -2 -30-3 -4 -35 2 3 4 Azimuth [deg] 60

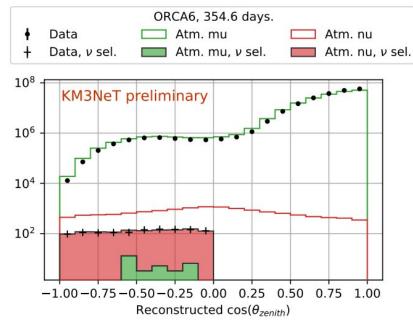


#### **ORCA-6** Data: We see v-Oscillation !

- Neutrino Oscillation
  - 355 days of data with ORCA-6
  - good matching between data and simulation
  - good suppression of the atm. muons
  - no-oscillation hypothesis excluded at  $5.9\sigma$



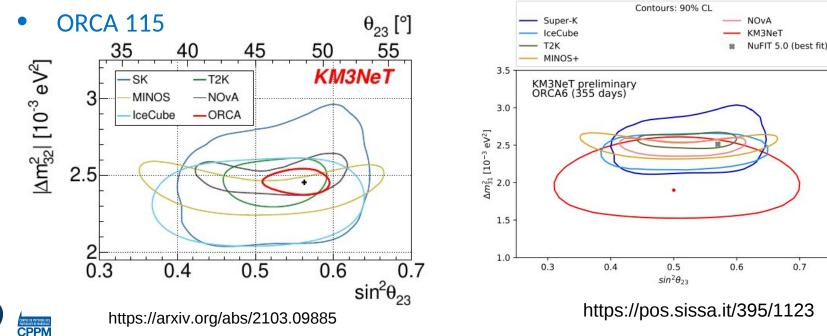




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#### We can measure v-Oscillation parameters !

- Atmospheric oscillations parameters contour
  - getting competitive with other experiment
- Prospects
  - new analysis with improved reconstruction (energy)
  - more data with more DUs

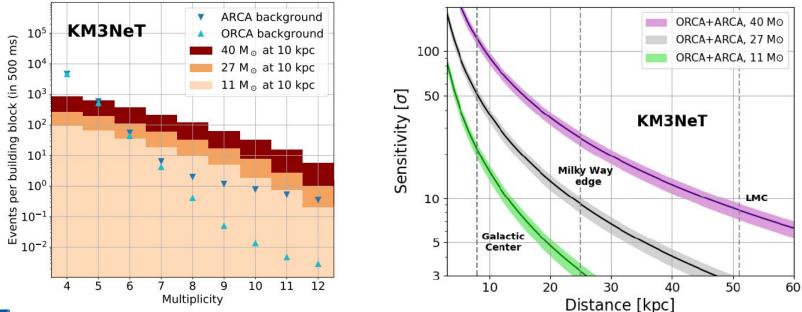


#### **Core Collapse Supernova**

• Core Collapse SuperNova (CCSN)

Eur. Phys. J. C 81, 445 (2021)

- O(10) MeV v detected from increased PMT coincidence rate
- CCSN analysis implemented real time since 2019 and deployed in SNEWS2.0
- with ORCA-6 and ARCA-6 the horizon for 11-27 M<sub>sun</sub> is 6-11kpc

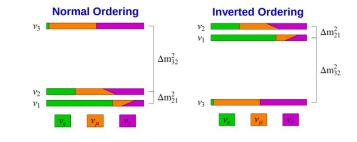


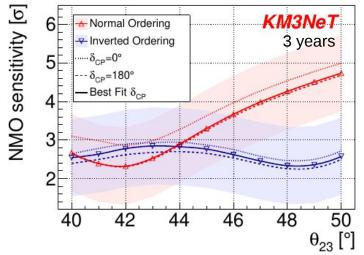


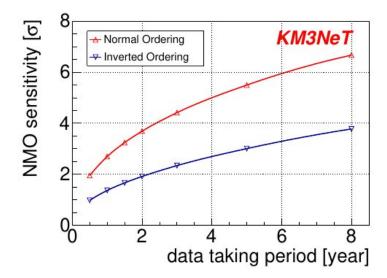
#### **Physics with Full ORCA**

#### https://arxiv.org/abs/2103.09885

- Update of the sensitivity to atmospheric oscillations
  - new trigger algorithm
  - new reconstruction method
  - additional systematic uncertainties



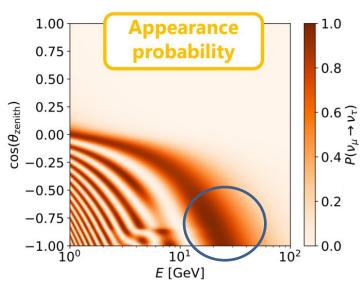


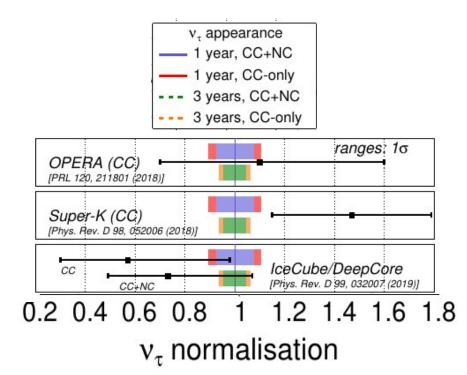




#### ντ appearance

- Few experiments sensitive to ντ:
  - ντ charge current interaction only possible if Ev > 3.5 GeV
  - dominant probability transition at 25 GeV
  - very important channel to constrain PMNS matrix unitarity (and indirectly sterile v





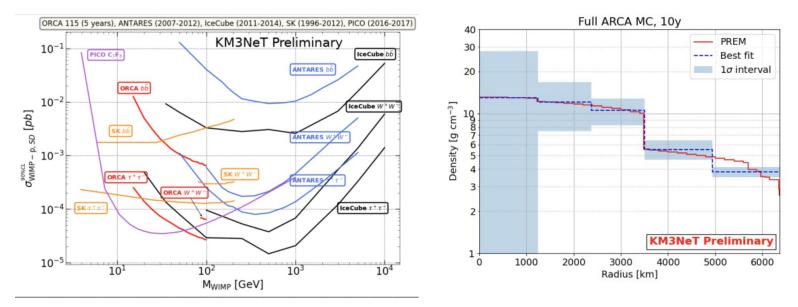
## **Combination with JUNO**

Baseline: 52.5km Oscillation parameters from NuFit5.0 (Table 1) Energy resolution:  $3 \% / \sqrt{(E)}$ Synergy between ORCA and JUNO No Oscillation 0.8 scillation NMO Oscillation IMO JUNO's precise determination of  $\Delta m^2_{31}$ Arbitrary Unit 6'0 boosts the ORCA NMO sensitivity  $\theta_{12}$ NMO at Joint paper being published: JUNO • θ<sub>13</sub> arXiv:2108.06293 0.2  $\Delta m_{2}^{2}$  $\Delta m_{22}^2$ 00 5 3 6 Visible Energy (MeV) True Normal Ordering (test IO) **True Normal Ordering** |<sup>∼</sup>×150 10 Asimov sensitivity [σ] COMB. 8 ORCA ORCA 100 COMB 6 4 50  $\Delta^{\text{JUNO}}_{\chi^2}$ JUNO  $\Delta_{\chi^2}^{\text{comb}}$ JUNO  $\overset{\text{ORCA}}{\Delta}\chi^2$ 2 0 0 -2.5 -2.7 -2.6 -2.4 -2.3 2 8 10 6 4  $\Delta m_{31}^2 [\cdot 10^{-3} \text{ eV}^2]$ Time [years] **CPPM** 

**CNrS** 

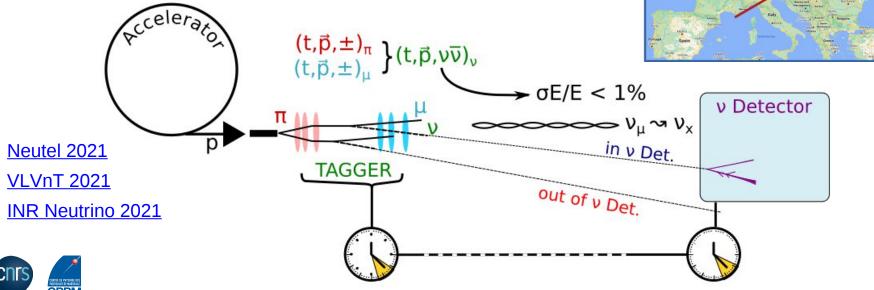
## **Others main physics topics**

- Several topics can also be addressed:
  - Indirect dark matter search from the Galactic Center and the Sun
  - **Earth tomography** via matter effect in the v oscillation / absorption of the v flux
  - v astronomy at intermediate energies GeV TeV from transient sources with ORCA:
    - choked GRB, Novae outburst, GW sources, flare of the Sun...
  - Combined ORCA/ARCA (ANTARES) analyses to increase the energy range with maximal sensitivities.



# Tagged P2O

- Neutrino beam from Protvino to ORCA
  - Re-use the same ORCA detector (no upgrade)
  - ORCA mass (6 Mton) allows to operate at lower beam intensity
  - Si detectors in beam line to track all  $\pi \pm$  and  $\mu \pm$
  - v produced by  $\pi \rightarrow \mu v$  can be presicely reconstructed: **tag-v**
  - v interacting in ORCA matched individually with tag v





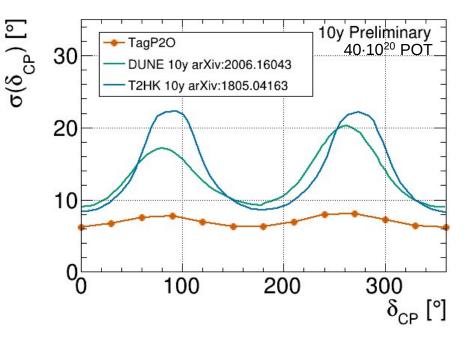
arXiv:1902.06083



#### **Tagged P2O Status**

#### • Tagging technique

- ANR obtained to study the technique
- Experimental validation ongoing at CERN-SPS with NA62: data being collected
- Beam line being studied @ CERN-PBC & IHEP Protvino
- Trip to IHEP Protvino in December
- Expected sensitivity
  - far detector: ORCA-115 with the perf. as for atmospheric v (conservative)
  - exploit excellent energy resolution (resolve osc. pattern, reduce NC bkg)
  - reduced systematic uncertainties (beam composition, energy scales)





#### Conclusions

- KM3NeT technology and analysis method are validated
  - 6 DU operational for more than 1.5 year
  - data collected already allows to observed ν-oscillations at 5.9σ
  - atmospheric osc. parameters are measured
- Construction is progressing
  - 7 more DUs ready to be deployed
  - 13 more DUs under construction for the next year
- Physics Case
  - strong opportunities for several world best/first measurements [NMO, Tau, Sterile]
  - **Tag-P2O**: a new avenue for v physics which allows
    - unique opportunities (e.g. precise measurement of  $\Delta_{CP}$ )
    - to attract new collaborators in KM3NeT
    - a full exploitation in the longer term of money/efforts invested in ORCA

